

# SuperDARNのRBSP(VAP)/ERGモードの 運用について

## SuperDARN RBSP(VAP)/ERG collaboration and operation modes



A. S. Yukimatu (NIPR),  
N. Nishitani (STEL, Nagoya-U),  
T. Nagatsuma (NICT)

# Super Dual Auroral Radar Network (SuperDARN)

+Adak E&W-201209!

01:11:00 UT 30 Jun 2011

01:11:00 UT

Hokkaido  
(+Hokkaido-W!)  
+Siberia

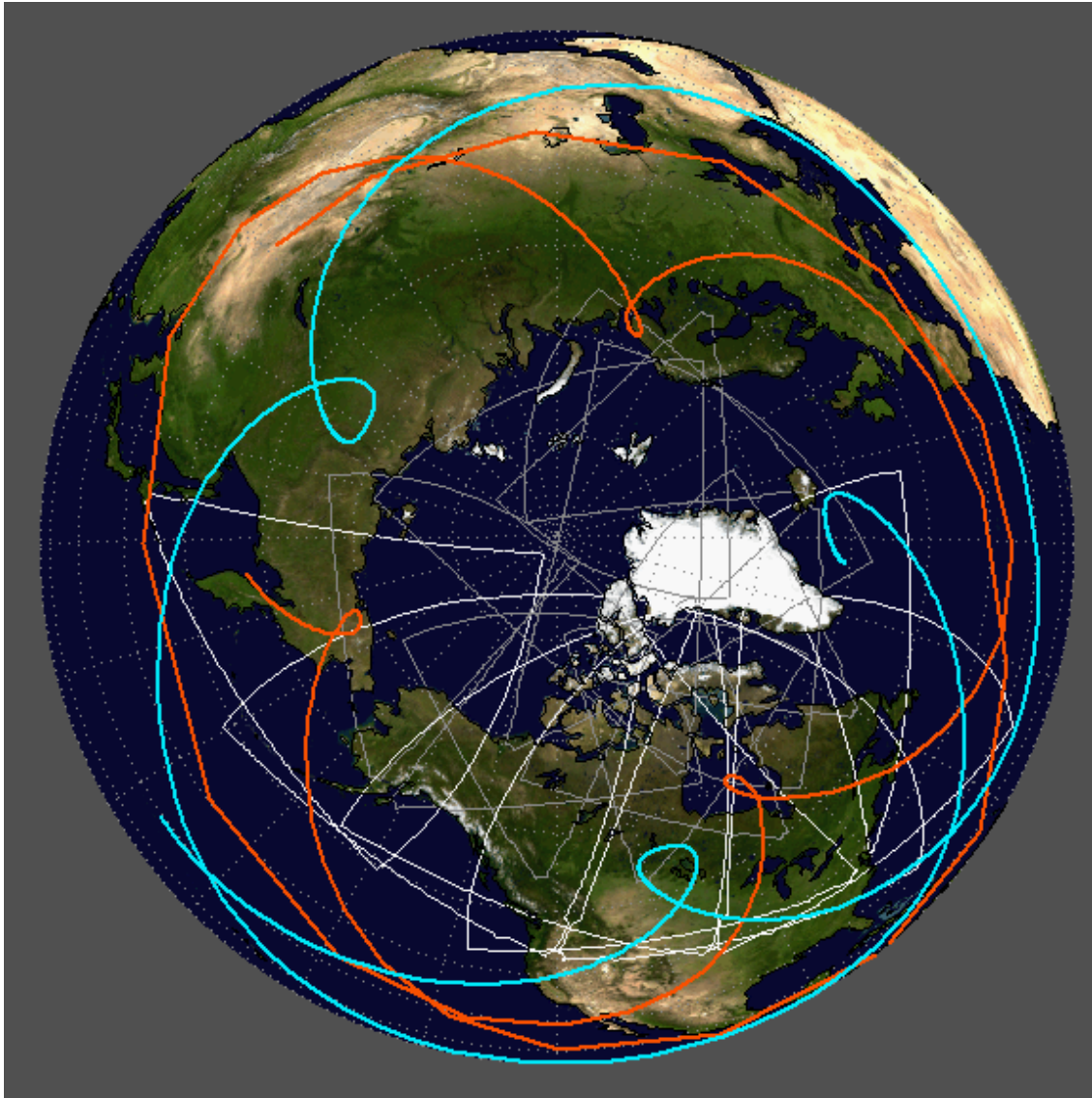
King Salmon

+South Pole-201301!

2 SENSU  
Syowa radars

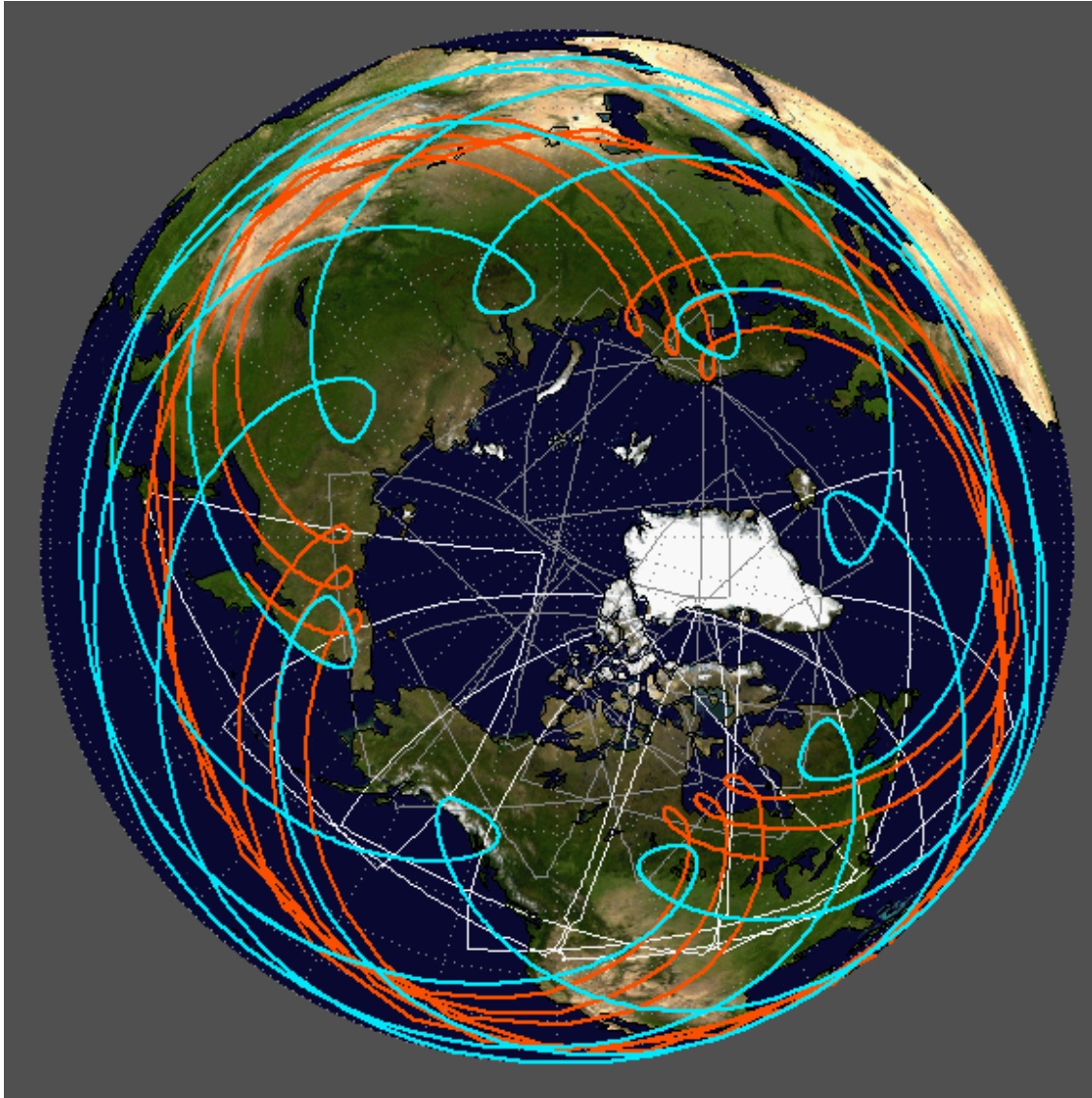
Number of operating HF radars: 28 (18 in the northern and 10 in the southern hemispheres) as of **June 30, 2011**

# ERG and RBSP(VAP) footprints and SD FOVs



ERG  
RBSP-A  
for 24 hour

# ERG and RBSP(VAP) footprints and SD FOVs



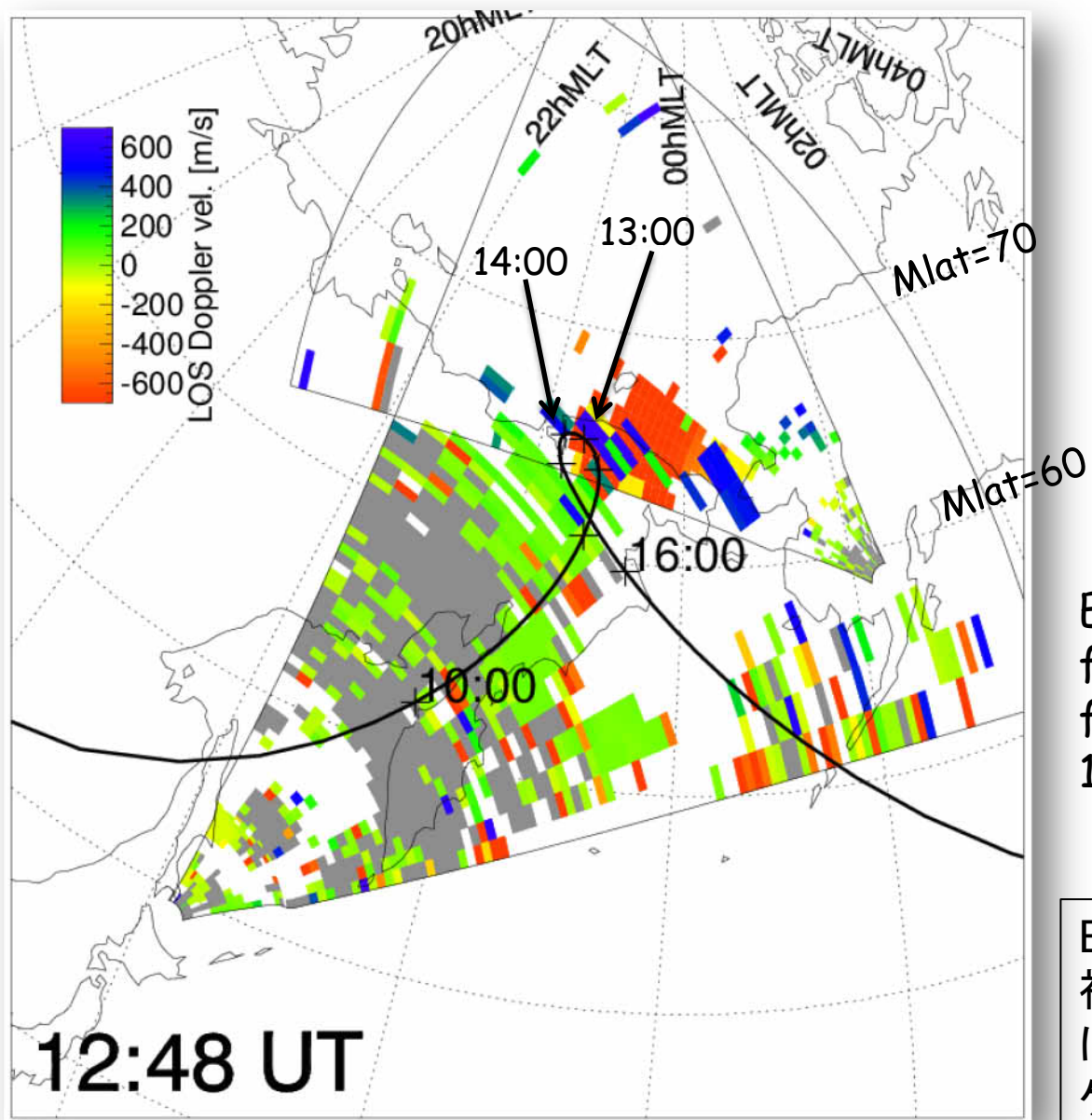
ERG  
RBSP-A  
for 72 hour

Courtesy of T. Hori

SuperDARN  
Hokkaido &  
King Salmon  
for [Jul. 17, 2007](#)  
12:48 UT

ERG satellite  
footprint  
for [Jul. 17, 2016](#)  
10:00-16:00 UT

ERGのfootprintがSD radarの  
視野中の定点(near apogee)  
に長期間(約5時間)滞在する  
ケース - 弱いsubstorm中の  
SAPS(西向高速flow)観測例



# SD Satellite collaborations

- As an important ground-based tool covering considerable portions of high to mid- latitude ionospheres at both hemispheres, SuperDARN has extensively collaborated with satellite missions like Geotail, CLUSTER, IMAGE, THEMIS etc.
- RBSP was launched to investigate IM recently (Aug. 2012).
- SD has been expanding its FOVs to sub-auroral and mid latitudes (StormDARN) as well as very high latitude (PolarDARN).
- SD PI group has decided to collaborate with **RBSP** mission.
- SuperDARN did also express the collaboration with **ERG** mission.
- Since last summer, discussions about **how to operate SD in conjunction with RBSP** have started and test operation started since last autumn.
- PI agreement has been updated so that **CT data** are now **immediately open to the public**. (also, CDF data for TDAS...)

# SD new PI agreement

- More clearly stated about “Open data policy”
- 5.3 Data Usage
- The SuperDARN Executive Council aims to operate a completely open data use policy for all data and will actively encourage the wider usage of the data by the scientific community. The “Rules of the Road” as outlined in Appendix 9 describe the policy in detail. Provision may be limited by available resources and subject to reasonable specific requests and variable lead times for the release of data depending on the operating time category (see Section 4) as follows:
  - (i) Common Time data
  - Common Time data are immediately available to all once data have been made available to all Principal Investigators.
  - Level 2 data products based on level 1 Common Time data can also be made available to all on the same time scale provided these data are produced by software approved by the SuperDARN Executive Council. A current example of a level 2 data product is the SuperDARN Map Potential. All Level 2 data products currently approved by the SuperDARN Executive Council are listed in Appendix 6. All plots and numerical data produced must be accompanied by a “Rules of the Road” notification outlined in Appendix 9.
  - (ii) Special Time data.
  - Access to Special Time data is restricted for a period of **one year** following the distribution of the data. During this time the data are for the exclusive use of the SuperDARN Principal Investigators whose radars operated in the Special time mode and Co-Investigators, Guest Investigators and Users designated by these Principal Investigators. After one year access rights are the same as for Common Time data. Exclusivity may be waived by the relevant Special Time investigator before this.
  - (iii) Discretionary Time data
  - (iv) Higher level Data Products

# SD new PI agreement

- Appendix 9
- SuperDARN “**Rules of the Road**” for Data Usage
- **SuperDARN has an open data use policy – prior permission to access and analyse the data is not required.** *However,* the data user is strongly encouraged to establish early contact with any Principal Investigator whose data are involved in the project to discuss the intended usage. Data are often subject to limitations which are not immediately evident to new users. In addition, some data is embargoed for use by designated Investigators for a period of one year. SuperDARN and the organizations that contributed data must be acknowledged in all reports and publications (see Appendix 10). The SuperDARN Executive Council must be notified before data is redistributed through another data base. If you have any questions about appropriate use of these data, contact any SuperDARN Principal Investigator.



# SD RBSP mode

- SD may provide
  - **global convection maps** in quasi real time including mid- to high latitude (both hemi.) with typically 1-2 min temporal resolution
  - **ULF wave activities** over large MLT coverage from mid- to high latitude (both hemi.)
  - possibly higher spatial/temporal resolution data around satellite footprints
  - **any other requests/requirement?!**

# SuperDARN



originally designed  
to obtain ...

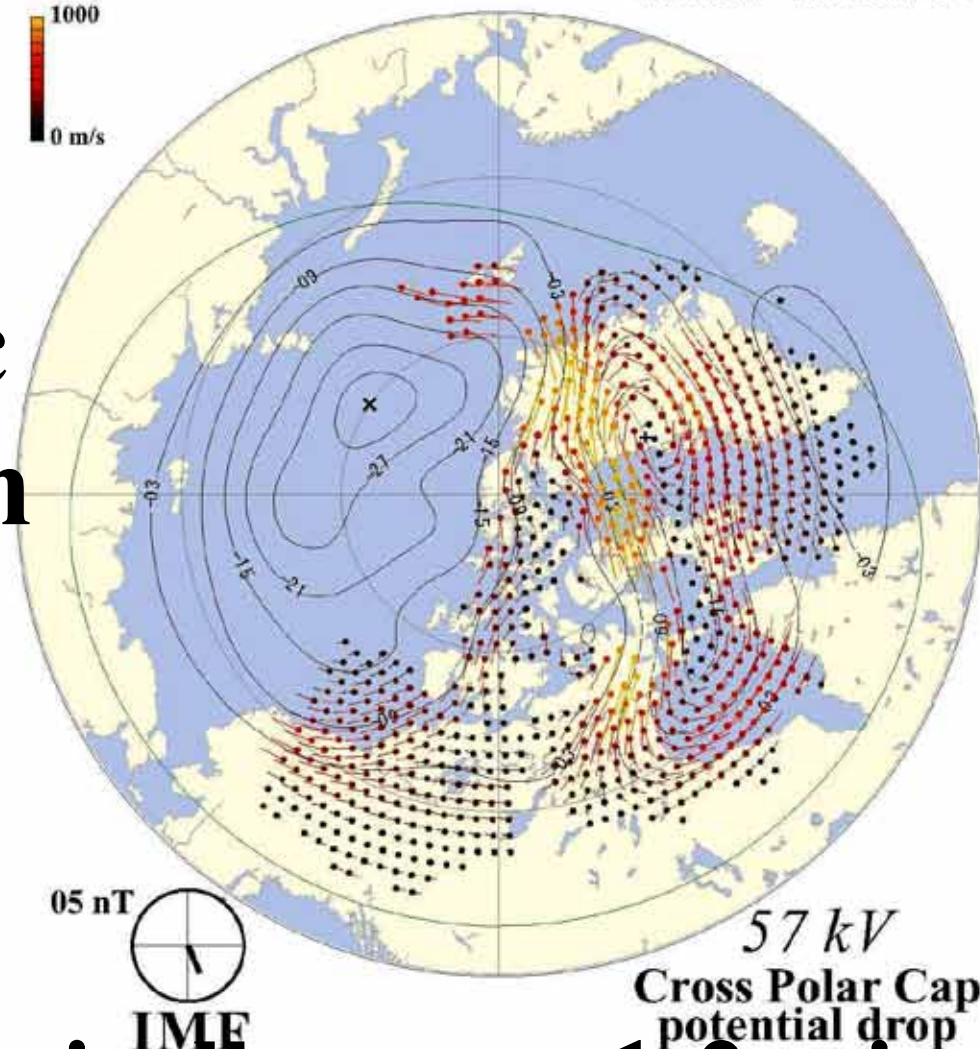
**global  
polar ionospheric  
plasma convection**  
*(...Ray's dream...)*

but as time goes by and SD grows,  
it has been realized that SD  
can address many more  
issues/scientific questions

## APL real time convection map

16 Oct 2001

08:40:00 ~ 08:42:00 UT

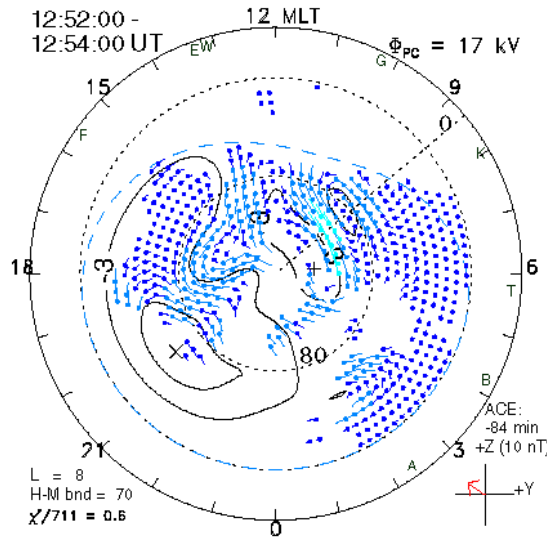
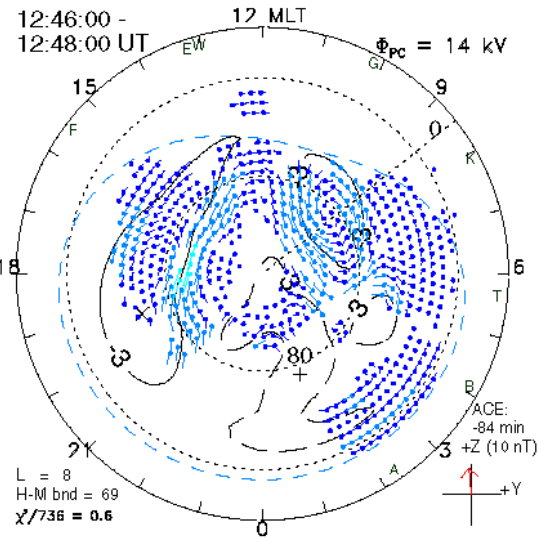


**typically every 1-2 min**

# dynamics of global convection

especially when IMF  $B_z$ ,  $B_y$  drastically changes

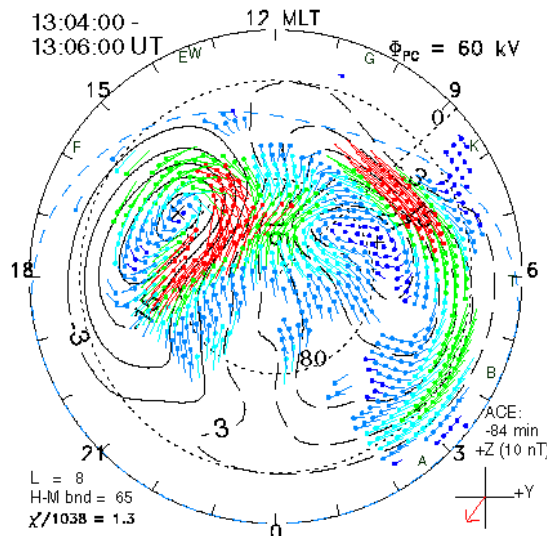
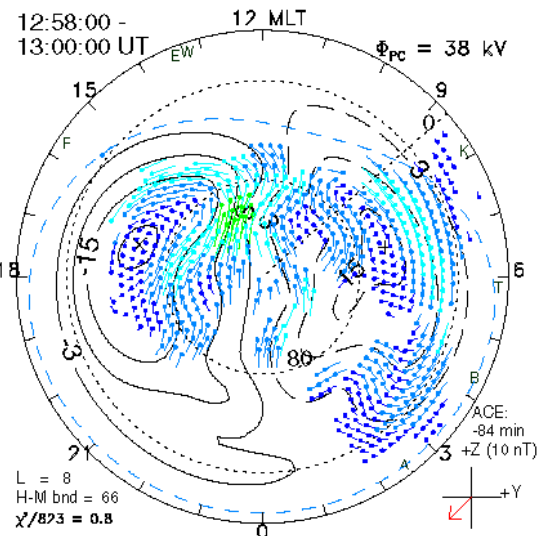
12/03/2001



- ~12:56: IMF  $B_z > 0$  stable condition
- 4 cell pattern @ 12:46
- 3 cell pattern @ 12:52
- 12:56: abrupt  $B_z < 0$  change happened!
- ~2min later aft  $B_z$  change:

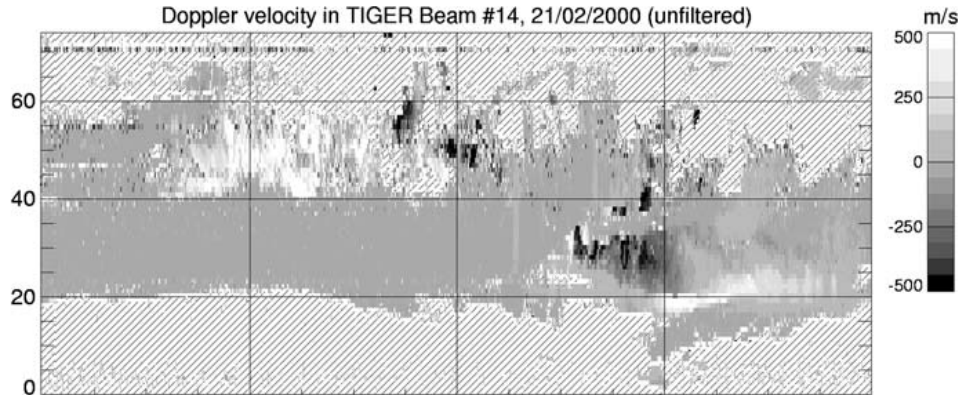
reconfiguration from  $B_z > 0$  pattern to  $B_z < 0$  pattern within 2-4 min.

● ~10 min later: enhance & strengthen  $B_z < 0$  pattern

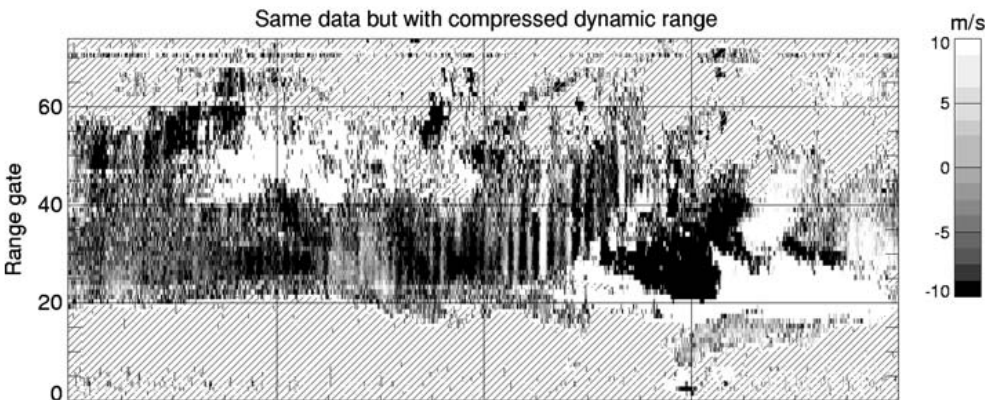


# SD MHD waves monitoring

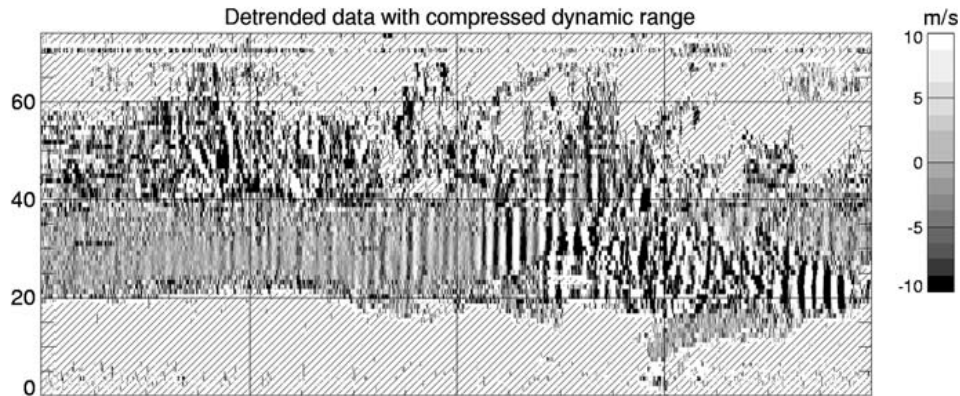
Doppler velocity in TIGER Beam #14, 21/02/2000 (unfiltered)



Same data but with compressed dynamic range



Detrended data with compressed dynamic range



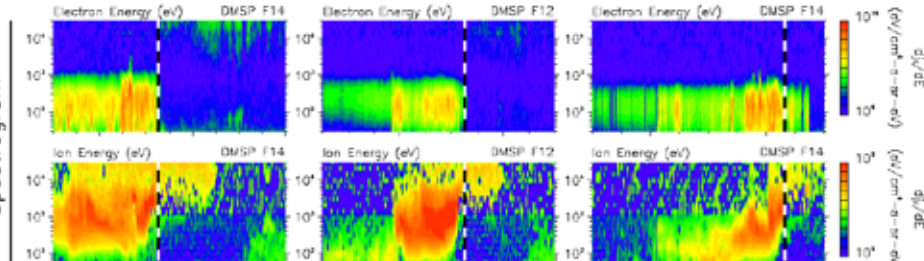
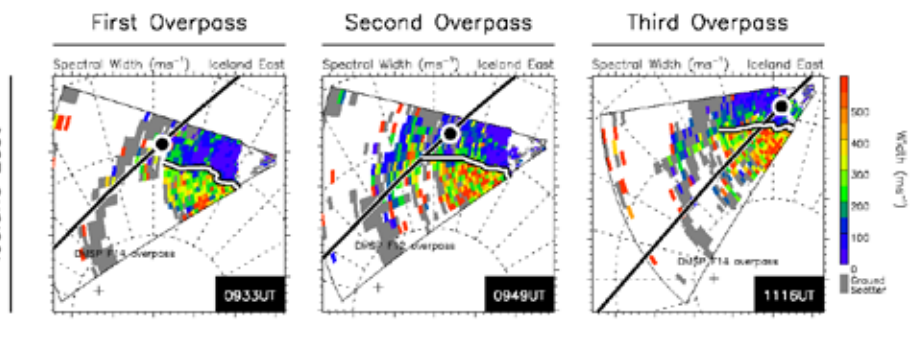
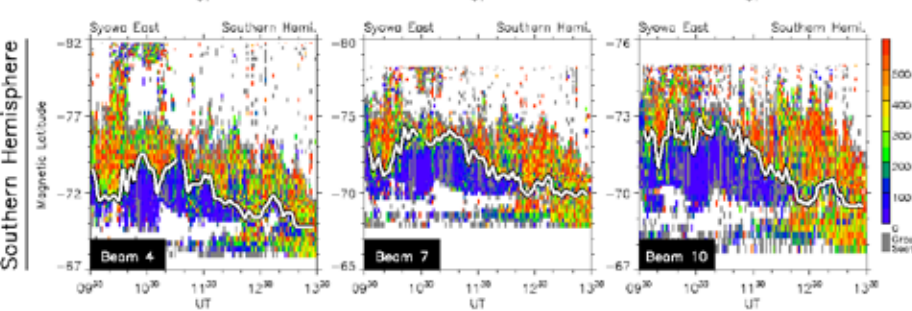
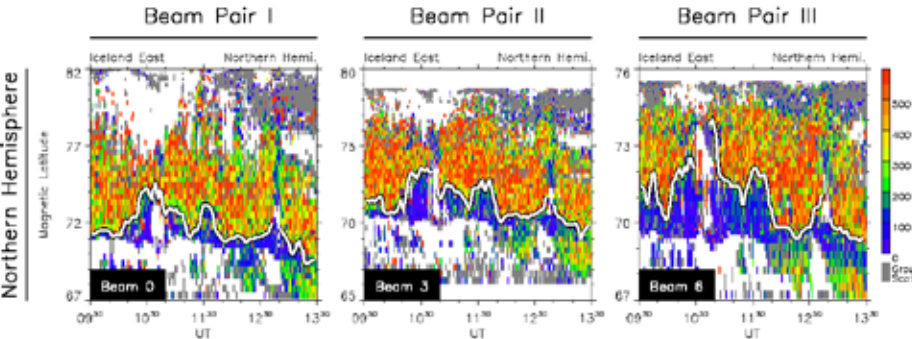
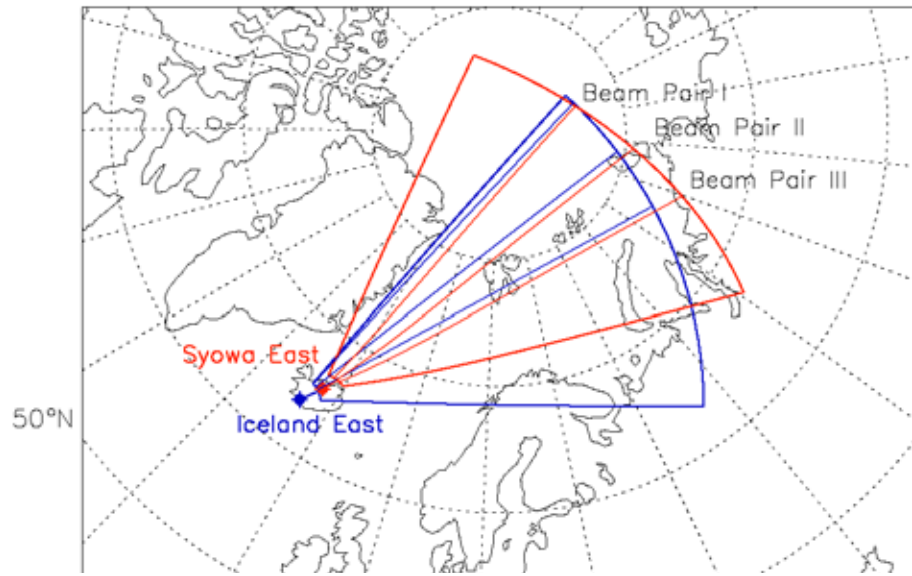
- Pc3~Pc5 ULF waves detectable w/ SD
- ULF waves also found in ground scatter echoes
- High-m ULF detectable
- 2-D and large MLT-Latitudinal distribution of ULF activities can be monitored

# SWB/OCB/FTE/Reconnection

## Spectral Width Boundary(SWB) as proxy of OCB?

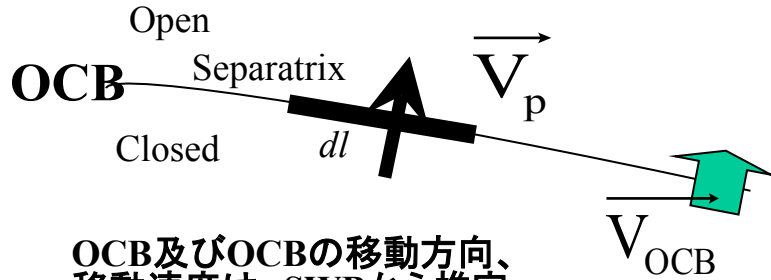
- Doppler Spectral Width sharply changes around Convection Flow Reversal (polar: higher width)
- Can Spectral Width Boundary (SWB) be a proxy of OCB (open/closed field line boundary)?

Hosokawa et al., Ann. Geophys., 2003



Hosokawa et al., Ann. Geophys., 2003

# estimate of Reconnection Rate



OCB及びOCBの移動方向、  
移動速度は、SWBから推定。

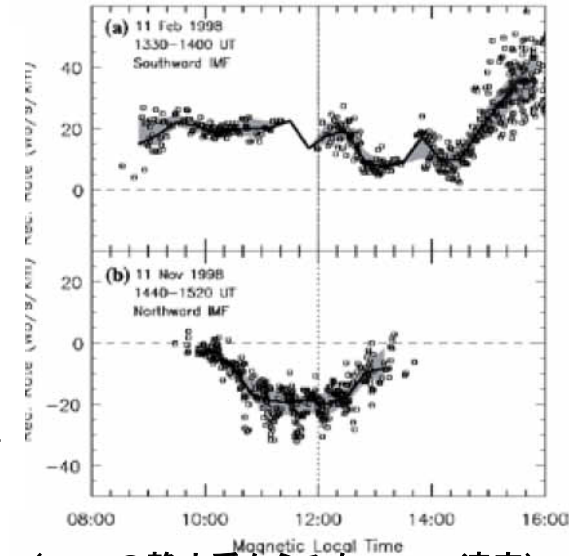
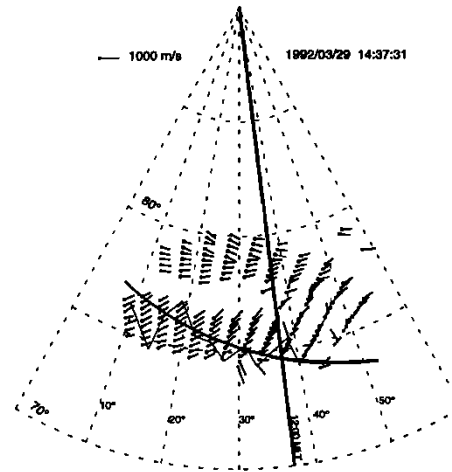
OCB上長さdlの境界線を超えて  
単位時間あたりに輸送される磁束Fは、

$$\frac{dF}{dt} = \vec{B} \times \vec{V} \cdot \vec{dl}$$

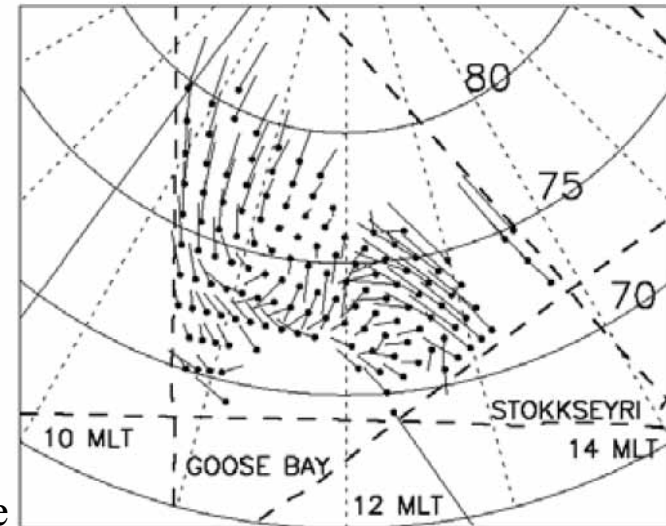
$$\vec{V} = \vec{V}_p - \vec{V}_{OCB}$$

## Testing Reconnection theories

- antiparallel reconnection hypothesis  
places where antiparallel happen
- component reconnection hypothesis  
subsolar point  
high sheath B strength & sheath flow stagnate

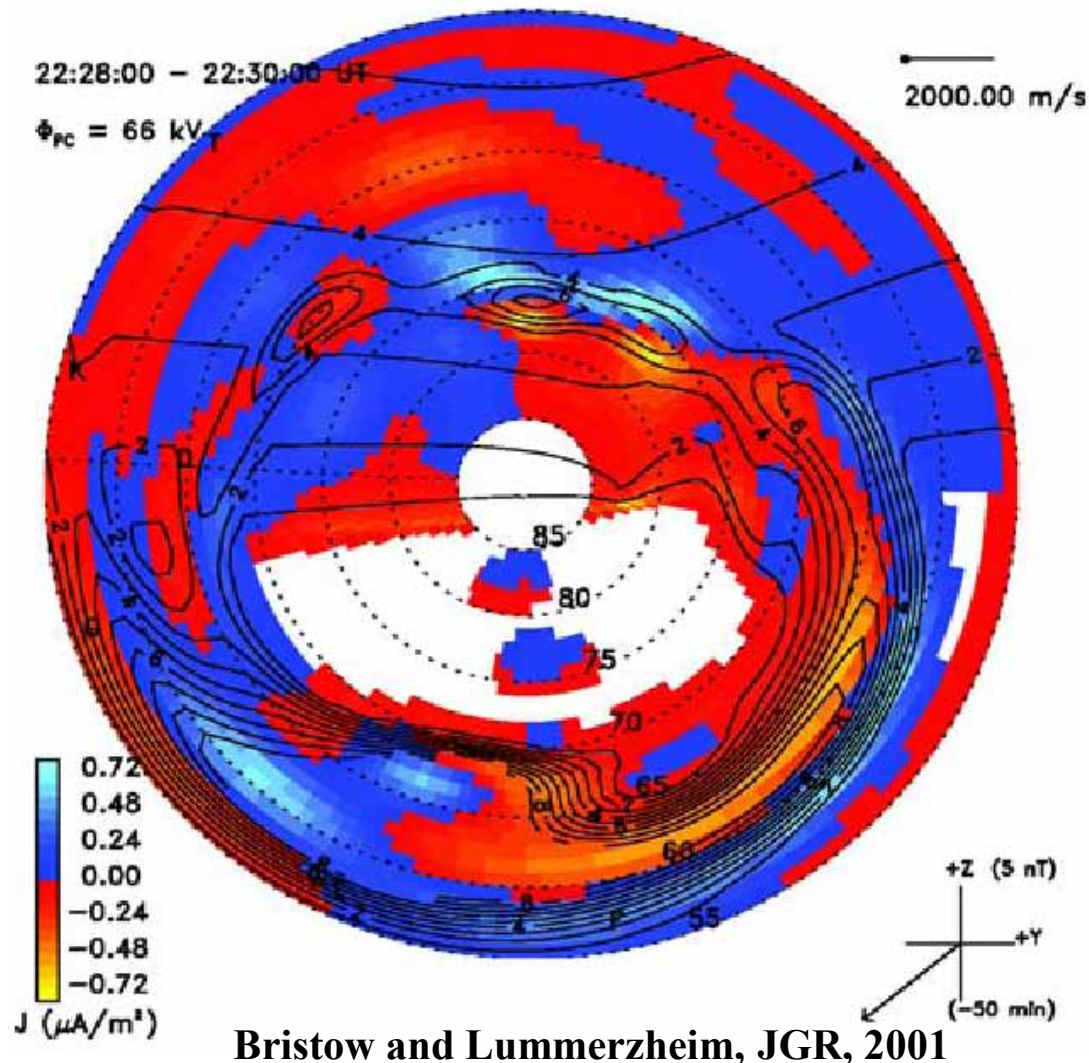


(OCBの静止系からみたplasma速度)



Baker et al., JGR, 1997  
Pinnock et al., Ann. Geophys., 2003,  
Chisham et al., Ann. Geophys., 2004b  
Coleman et al., JGR, 2001

# deduction of global FAC system



- From Polar UVI image auroral luminosity, particles precipitation into ionosphere can be inferred, which can then be combined with model for photo-ionization etc and Pedersen, Hall conductivity ( $\Sigma_P$ ,  $\Sigma_H$ ) at whole polar region can be deduced. (possible problems on reliability at low luminosity regions)

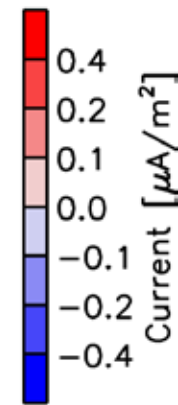
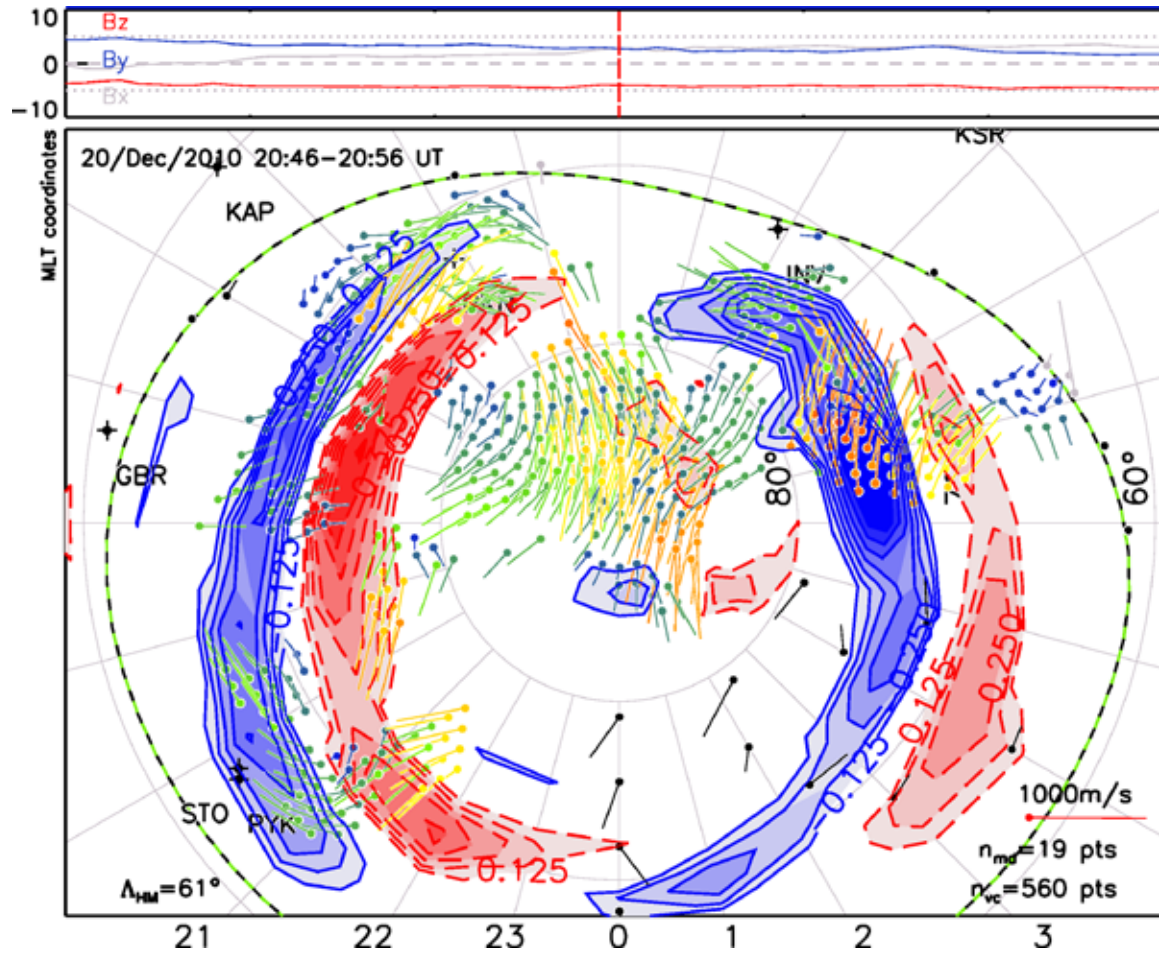
- By combining the deduced  $\Sigma_P$ ,  $\Sigma_H$  & SD 2-D global E fields,  $J_{\perp}$  can be obtained and FAC,  $J_{\parallel}$  can then be inferred (by  $\nabla \cdot J_{\perp}$ ).

- up(red)/down(blue) ward FAC, &  $\Sigma_P$  (contour)

- extensive validation with satellite magnetometer data etc required

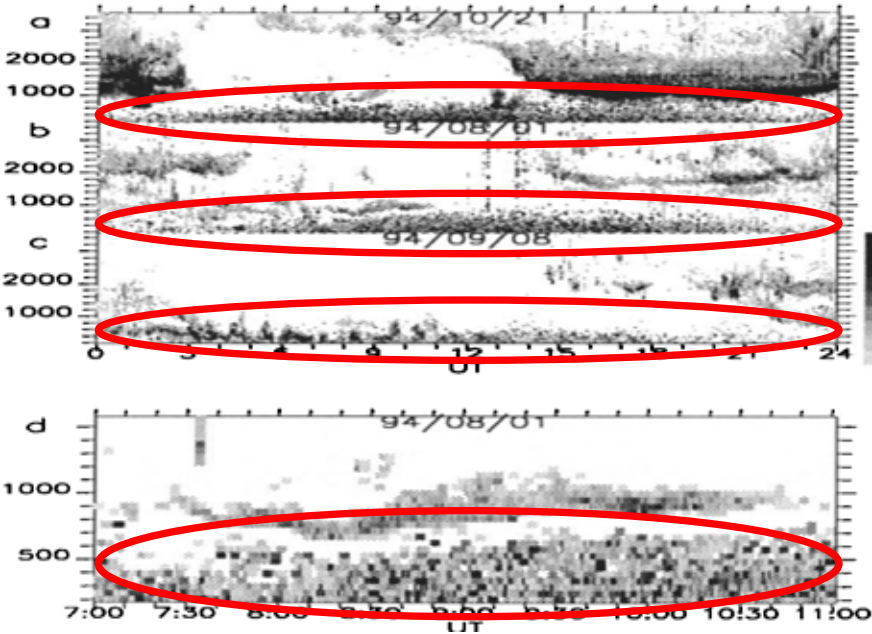
- 昼間側は太陽紫外線の影響がdominantで $\nabla \Sigma$ は小さく、 $\nabla \times V$ によるFACの評価でもよいが、夜側は、aurora活動による $\nabla \Sigma$ が大きく、 $\nabla \times V$ からFACの評価は困難であることもわかる。

# SuperDARN & AMPERE

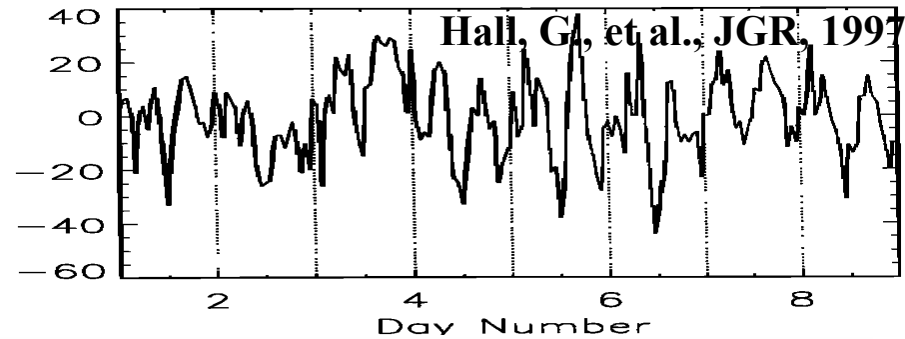




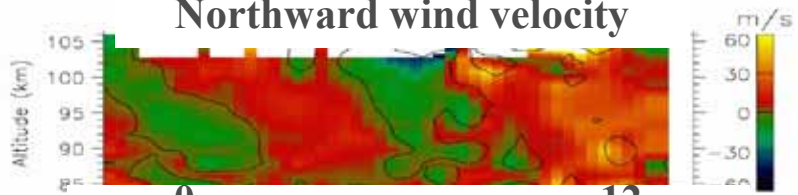
# neutral winds by SD meteor echoes



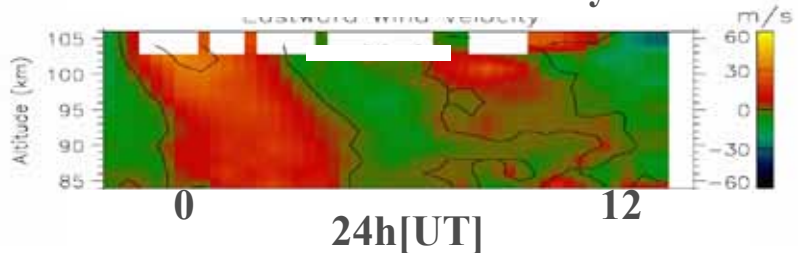
- GNREs (Grainy Near-Range Echoes) mostly found @  $r < 400$  km random echo power distrib. both in range & in time
- not depend on geomagnetic activity periodic appearance mostly every day



Northward wind velocity

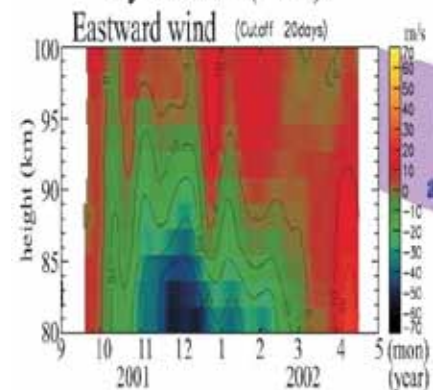


Eastward wind velocity

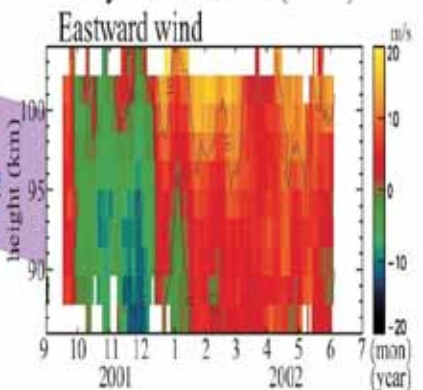


Downward phase propagating semi-diurnal tide

Syowa MF (FCA)



Syowa East HF (TMS)



basically good agreement!

Comparison with MF radar obs. - Good Agreement

● SD  $\Rightarrow$  meteor radar network!

# more topics related to mid-lat SD

- SC/Storm – Pc5 etc.
- Overshielding effect
- Disturbance dynamo theory – corotation lag...
- Region identification?
- ...

# SD scheduling and RBSP

- 3 categories

CT (Common Time)                     $\geq 50\%$  of each month  
getting 1-min global convection pattern  
(THEMIS mode – terminated this year)

DT (Discretionary Time)     $\leq 30\%$

ST (Special Time)                     $\leq 20\%$

- Proposed 2 months before each target month,  
final version fixed 1 month before the month
- Difficult to allocate storm time special mode
- **Newly introduce “Trigger-mode” for RBSP**

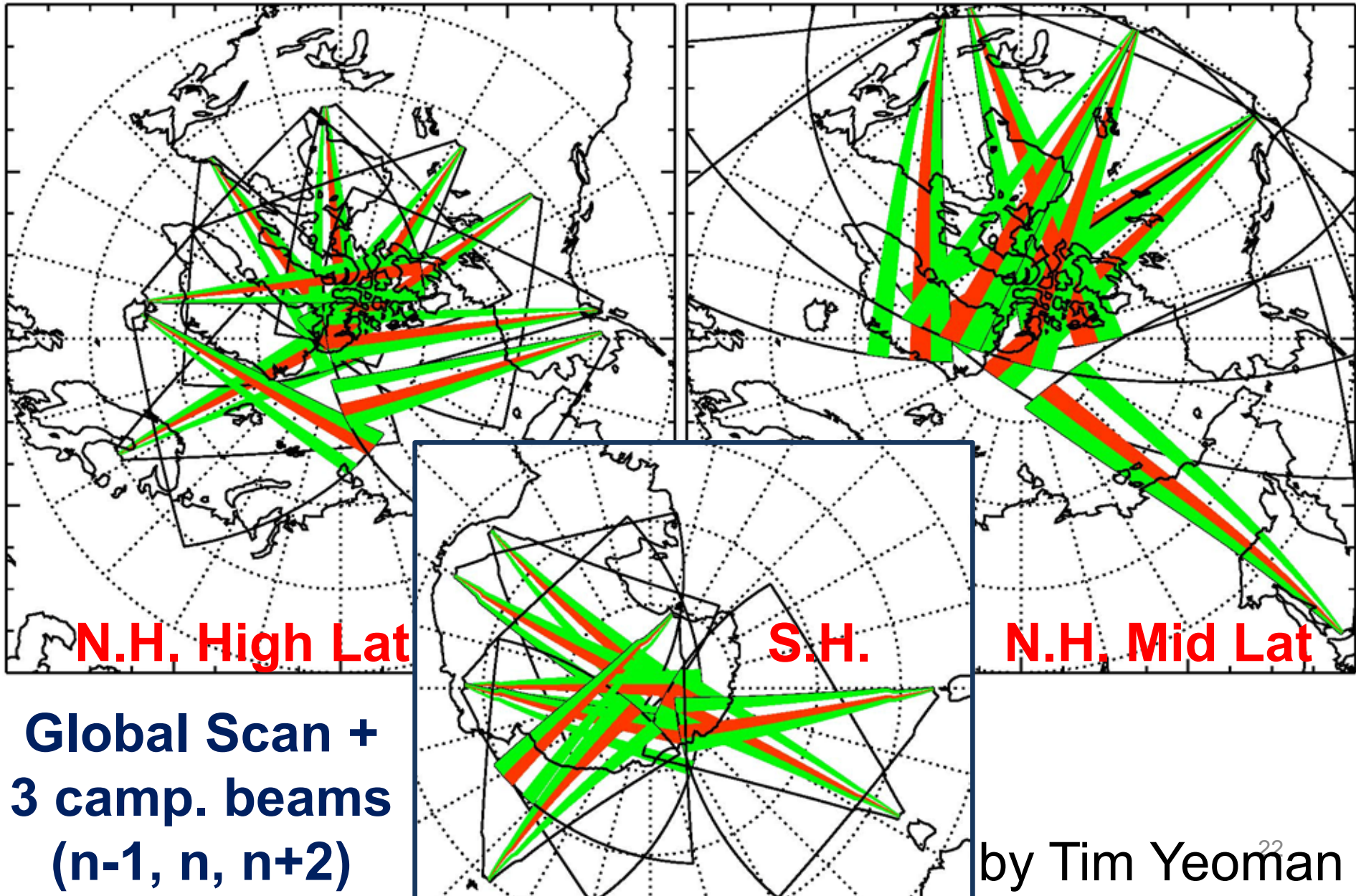
# SD RBSP mode (test)

- **CT-TRIG** mode introduced  
When a storm happens, SD changes from the normal CT mode to RBSP special mode during CT
  - **CT-TRIG data will be open like CT data**
- **ST-APOG** mode introduced  
ST-APOG mode is scheduled in advance for period when RBSP is near apogee whose footprint is in one of SD FOVs
  - **ST program but CT CPID will be assigned so ST-APOG data will be open like CT data**
- Override  
CT-TRIG overrides ST-APOG  
Priority btw CT-TRIG and DT decided by each PI  
Priority btw ST and CT-TRIG decided by each PI
- Tested Oct – Dec, 2012, will be reviewed Jan 2013

# SD RBSP mode – CT-TRIG (plan)

- Currently proposed and agreed mode:
- Global scan with 3-beams mini-scan interleaved **for global convection patterns and ULF activities**
- Integration time for each beam: 3 sec  
Range resolution: 45km (or less?)
- **Global convection patterns every 2 min  
high temporal resol. 3-beam data every 18 sec.**
- **Beam sequence:**  
**0, S1, 1, S2, 2, S3, 3, S1, 4, S2, 5, S3, 6, S1, ...**
- 3 beam selection: N-1, N, N+2(or 3)  
where N: poleward beam if possible
- **Trigger algorithm** - many debate....

# SD RBSP mode – CT-TRIG (plan)



# Triggers to initiate RBSP mode

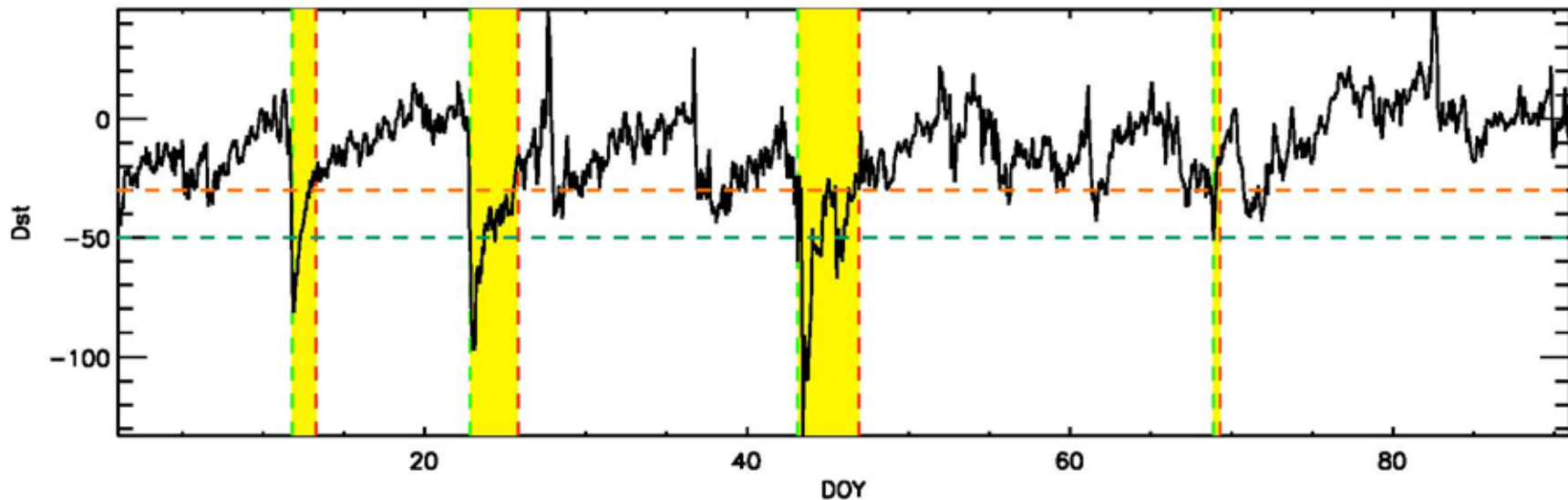
- Many debates...
- Current proposal
- **Simple Dst triggering algorithm**
- Started/triggered by  $Dst < -50nT$
- End countdown started by  $Dst > -30nT$
- Countdown time: 6 hours  
(terminated if  $Dst > -30nT$  for 6hours)
- **Avoid false positive (誤検出)**

# Simple Dst trigger algorithm

SUPERDARN PARAMETER PLOT  
2000010100 – 2000033123

Start Scan Mode

Stop Scan Mode



Dst Start Value = -50

Dst Start Countdown Value = -30

Countdown time(hours) = 6

4 runs

9.63409%

average duration: 52.0 hours

max duration: 91.0 hours

min duration: 10.0 hours

by Tim Yeoman

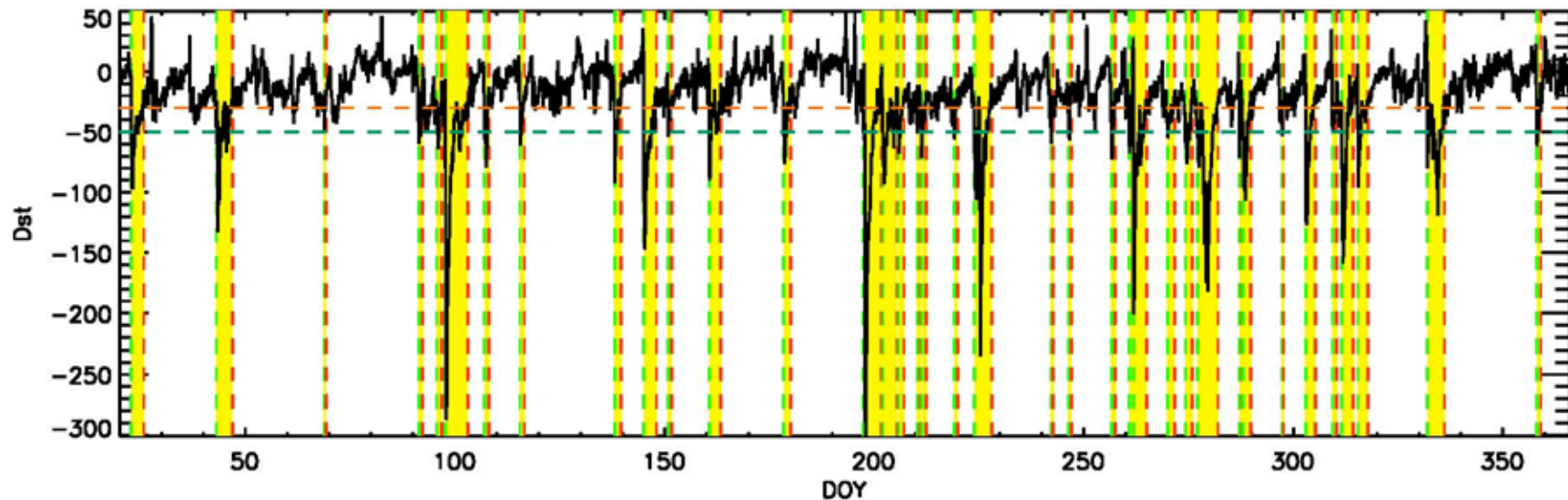


# Simple Dst trigger algorithm

SUPERDARN PARAMETER PLOT  
2000012000 - 2000123123

Start Scan Mode

Stop Scan Mode



Dst Start Value = -50

Dst Start Countdown Value = -30

Countdown time(hours) = 6

37 runs

20.8840%

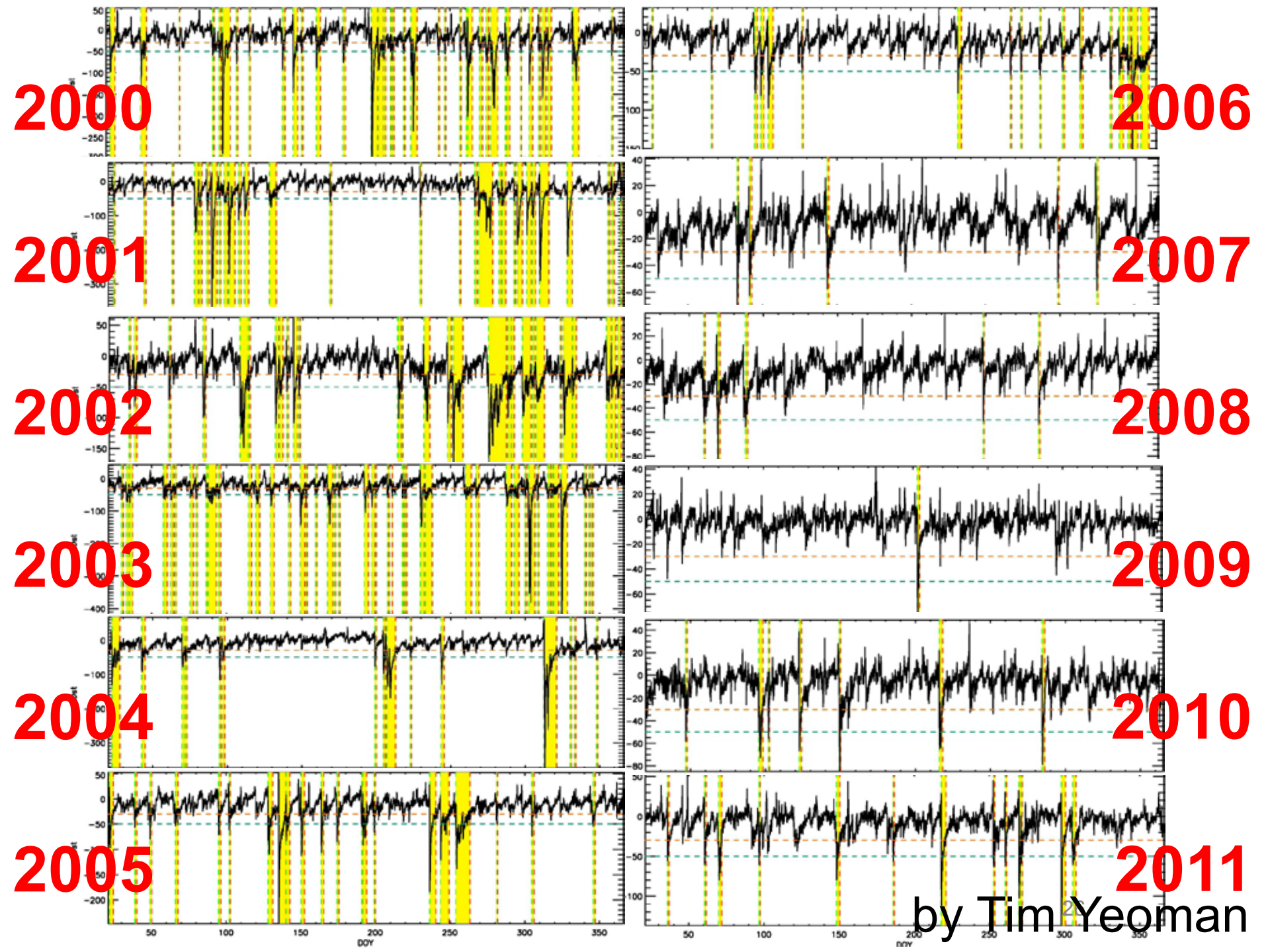
average duration: 46.9 hours

max duration: 127.0 hours

min duration: 9.0 hours

Ave sunspot number: 119.580

by Tim Yeoman

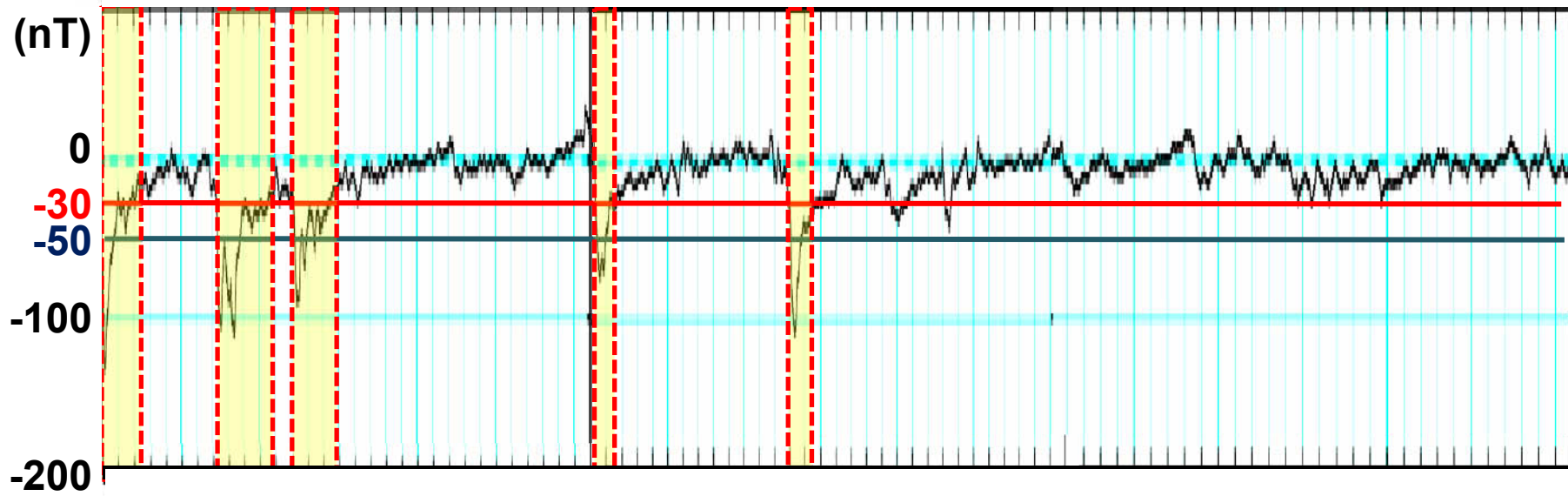


# Simple Dst trigger algorithm

SUPERDARN PARAMETER PLOT  
2012100100 - 2012123123

Start Scan Mode

Stop Scan Mode



Dst Start Value = -50

Dst Start Countdown Value = -30

Countdown time(hours) = 6

5 runs (19)

11.1865% (9.2808%)

average duration: 49.4 hours (42.8)

max duration: 82.0 hours (109.0)

min duration: 24.0 hours (15.0)

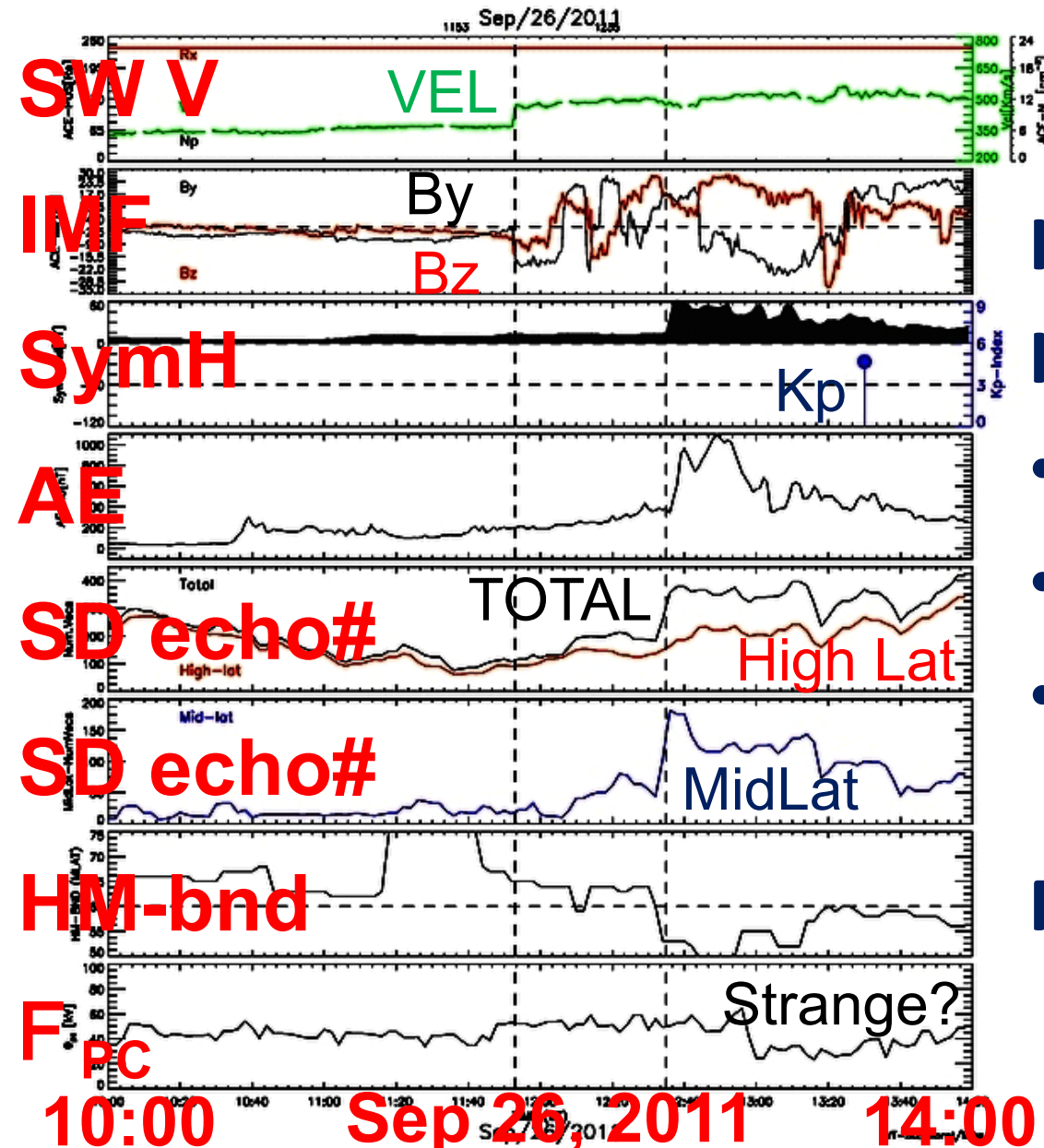
() means total or average in whole 2012

# Simple Dst trigger algorithm

Year	SSN	trigger time %	Max duration	Min duration	Mean duration	# of triggers
2000	120	21%	127	47	9	37
2001	110	20%	232	15	58	29
2002	104	24%	293	11	69	29
2003	64	25%	131	8	48	44
2004	40	10%	200	15	55	15
2005	30	15%	228	10	59	21
2006	15	9%	130	12	44	17
2007	8	1%	36	14	25	5
2008	3	2%	35	16	26	5
2009	3	0.4%	35	35	35	1
2010	16	3%	62	14	35	7
2011	56	6%	81	16	36	13
<b>2012</b>	<b>66</b>	<b>9%</b>	<b>109</b>	<b>15</b>	<b>43</b>	<b>19</b>

- Does not miss anything too significant
  - Does not run too often
  - No false positives (important)
  - Reasonable trigger condition to start with until something more sophisticated is developed.
- by Tim Yeoman(+Sessai)

# Dst Trigger mode – problems



However,

It does miss...

- Pre-onset
- Onset
- Early Development of the storm cycle

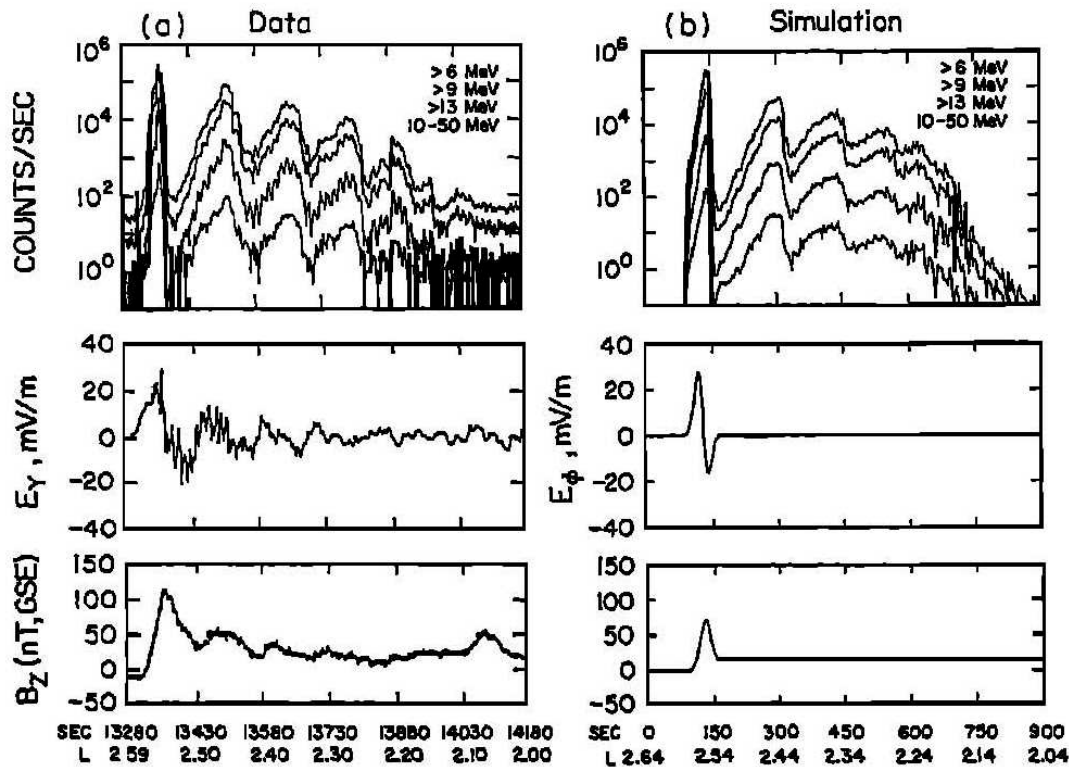
Prediction required...

by Mike Ruohoniemi

# Triggered by SCs?

- Some storms -no SC
  - depending on CME/CIR storms
- Miss SCs themselves!
  - possible drift echo event with large electric field...

Energetic Electron, Electric and Magnetic Fields  
24 March 1991



X. Li et al., GRL, 1993  
Miyoshi & Kataoka, 2005  
Kataoka & Miyoshi, 2007

# Triggers to initiate RBSP mode

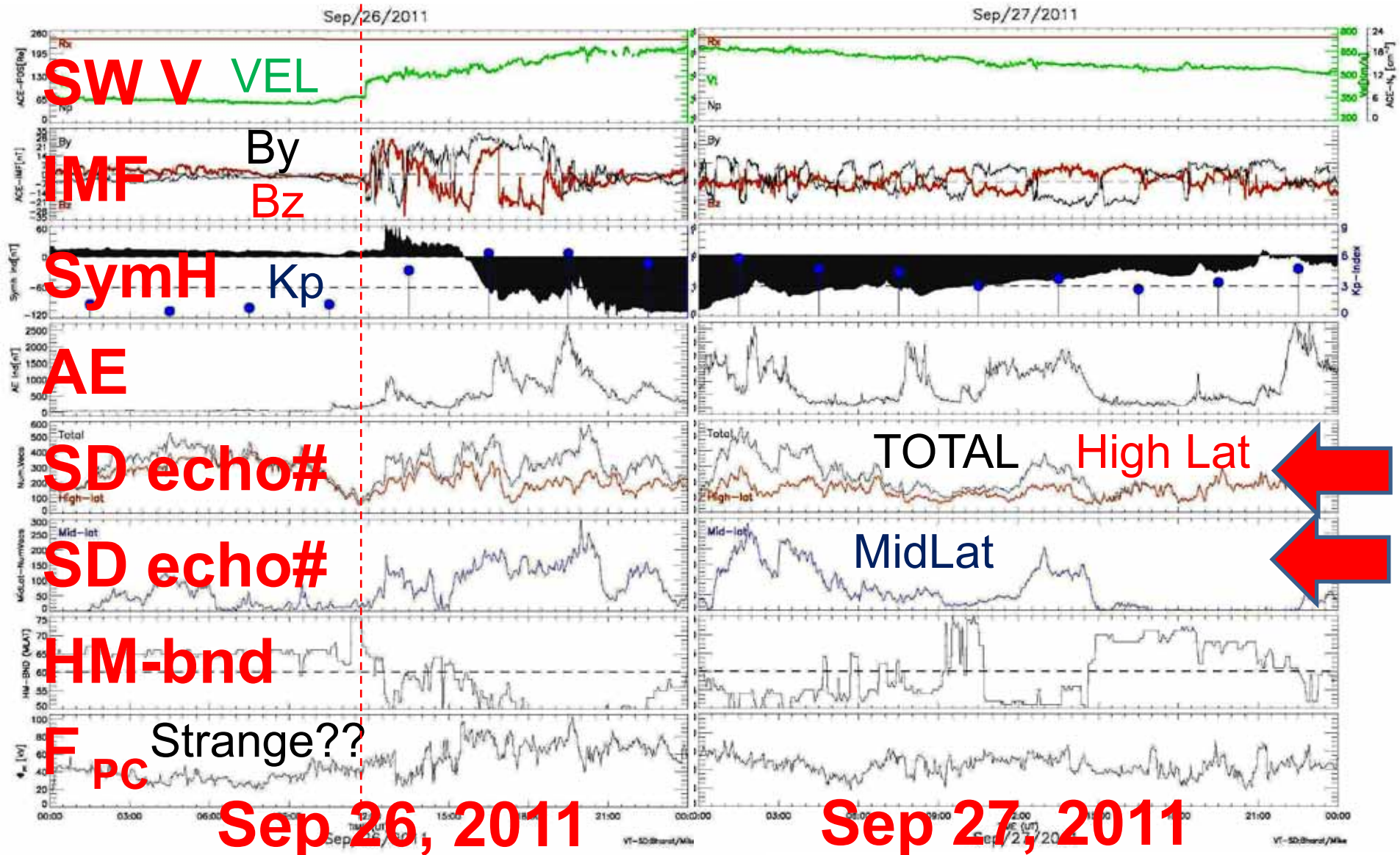
- Many arguments on how to trigger RBSP mode
- Triggered by **SCs**? – no SC CME/CIR, miss SCs!!!
- by **IMF (Bz) and Solar Wind parameters @ L1?**
- by **Simulation predictions?**
- by RBSP special campaign mode trigger if any?
- should always run RBSP during all CT time for longer monitoring including quiet period?!
- Some algorithm for cases of **false positives**
- Important scientific targets esp. space weather, too!  
Great if SD-Jpn/ERG group to propose for better ideas!

# SD ST-APOG mode

- A radar whose FOV has RBSP footprint during apogee periods (for several hours?):  
some sophisticated special mode?  
e.g., high spatial/temporal resolution mode?
- Other radars during ST-APOG period:  
run normal CT mode?  
run RBSP CT-TRIG mode?
- Scientific target?!?
- SD not fixed the details of this mode  
- **inputs will be welcome...**



# SD echoes during a Storm



# Summary

- SD-RBSP(VAP) collaboration is about to be started. SD also expressed support for ERG mission.
- Currently, SD's RBSP(VAP) mode is discussed in SD community.
- Triggering mechanism (how to trigger RBSP mode) is one of the important issues – and it is one of the important research targets, too.
- Sci. targets during ST-APOG mode - investigated
- Discussions on SD-RBSP mode will be valuable for planning of the future ERG-SD collaboration.