

#### SI-induced transient convective oscillation: comparison of SuperDARN observation and global MHD simulation

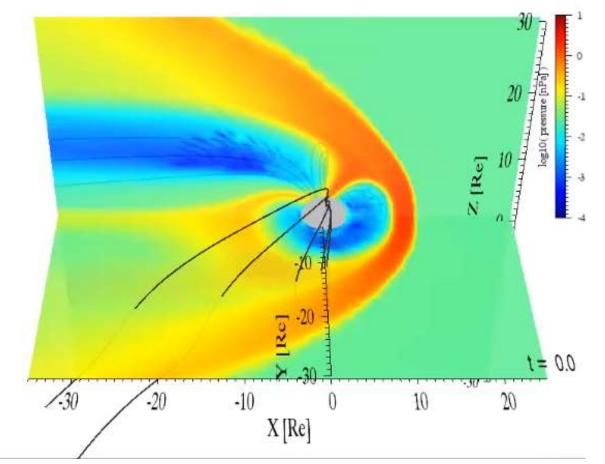
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# Today's talk

- 1. Introduction
- 2. Motivation & Objectives
- 3. Data & Method
- 4. Results
- ▶ 5. Discussion
- 6. Summary and Conclusion

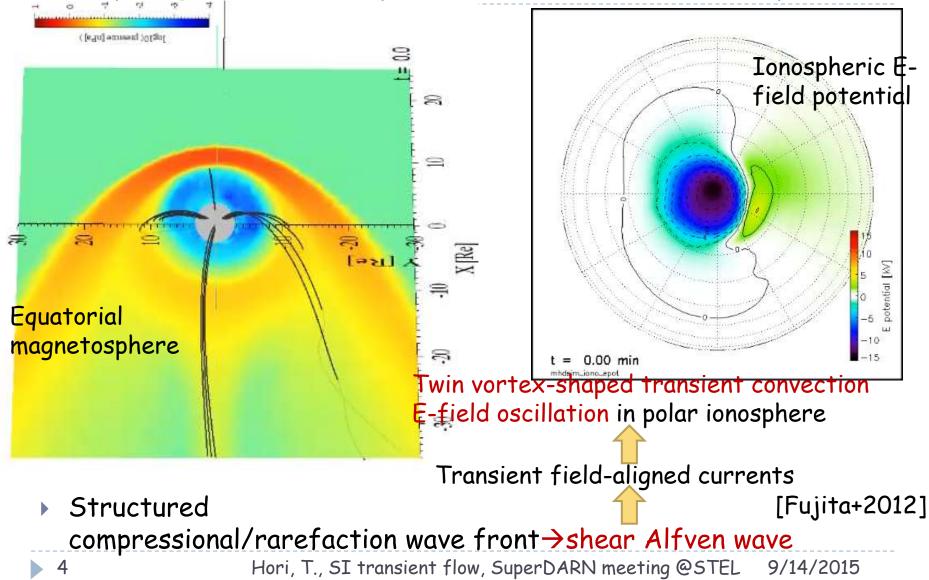
#### 1. Introduction Sudden compression/expansion of magnetosphere



Transient, but drastic changes in structure, E-M field are induced in the magnetosphere. Hori, T., SI transient flow, SuperDARN meeting @STEL 3

9/14/2015

# 1. Introduction: Coupling with dissipative, inductive ionosphere



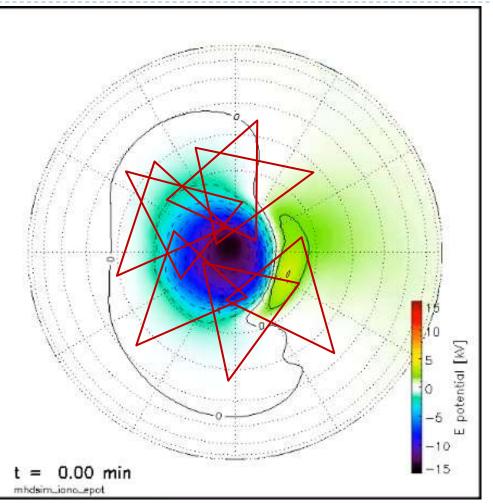
# 1. Introduction SI-assoc. flow perturbation by SuperDARN

Ionospheric E-field potential

ExB drift of ionospheric electrons

SD can measure its line-ofsight velocity component as a Doppler shift

We examine <u>"E-field"</u> evolution associated with SI



## 1. Introduction SuperDARN statistics for SI transient flow

Nothing but a mess if just

plotting all velocity data ...

1000

500

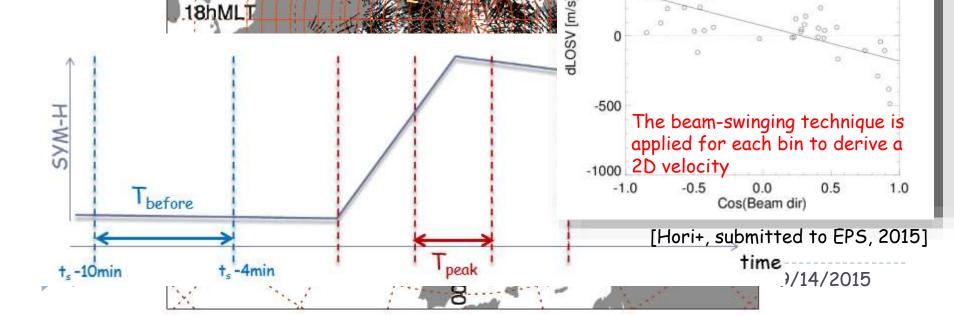
[e.g., Makarevich+2007]

All data, 73 < MLAT < 76, 16 < MLT < 18

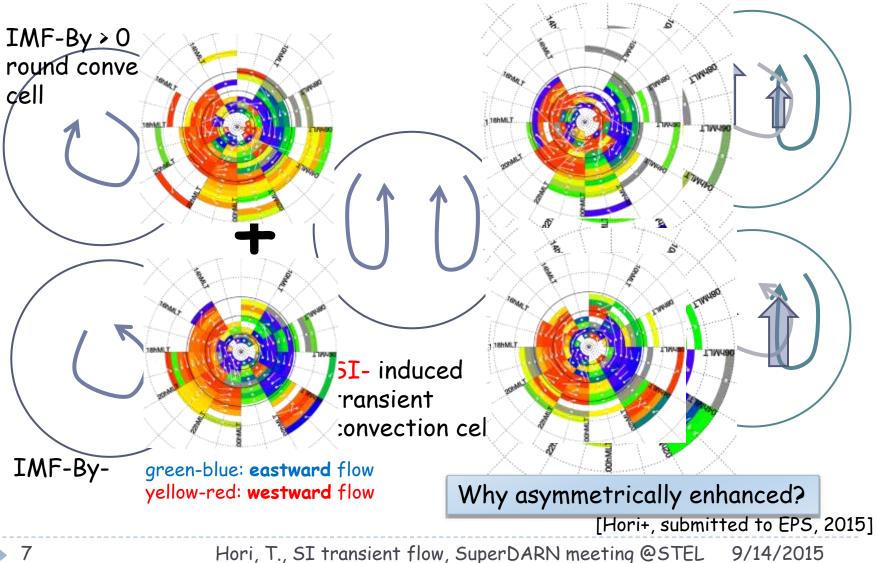
Vtrue [m/s]: 253.2

cos(beam dir) for dLOSV=0: 0.29 Vtrue angle [deg]: -16.9

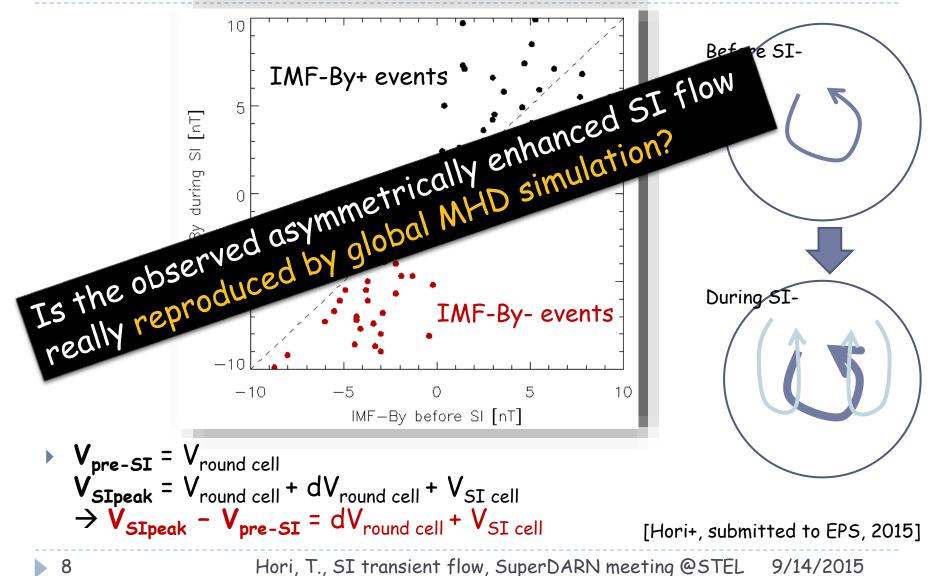
All N. hemis. SuperDARN Doppler velocity data for 192 SI+ and 179 SIevents during 2007early2014 were statistically analyzed to deduce a transient flow pattern.



# 1. Introduction pre-existing round cell and SI- cells: SD obs.



# 1. Introduction IMF-By intensification and round cell evolution

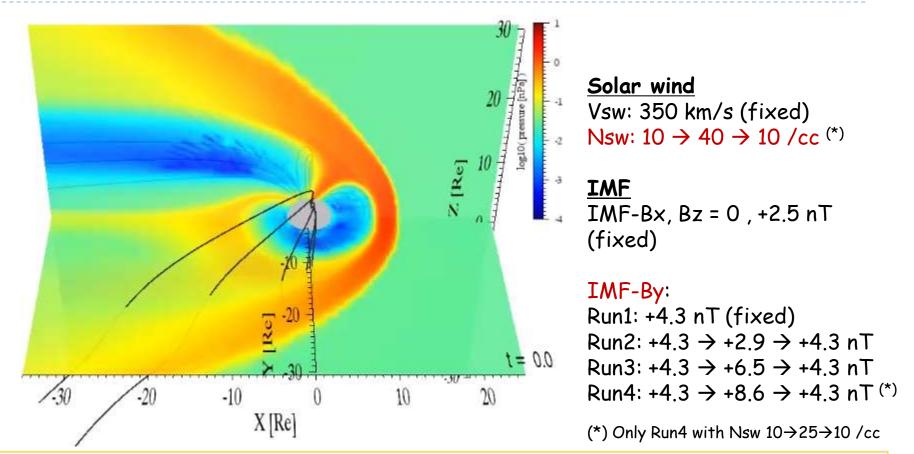


# 2. Motivation & Objectives

### Motivation

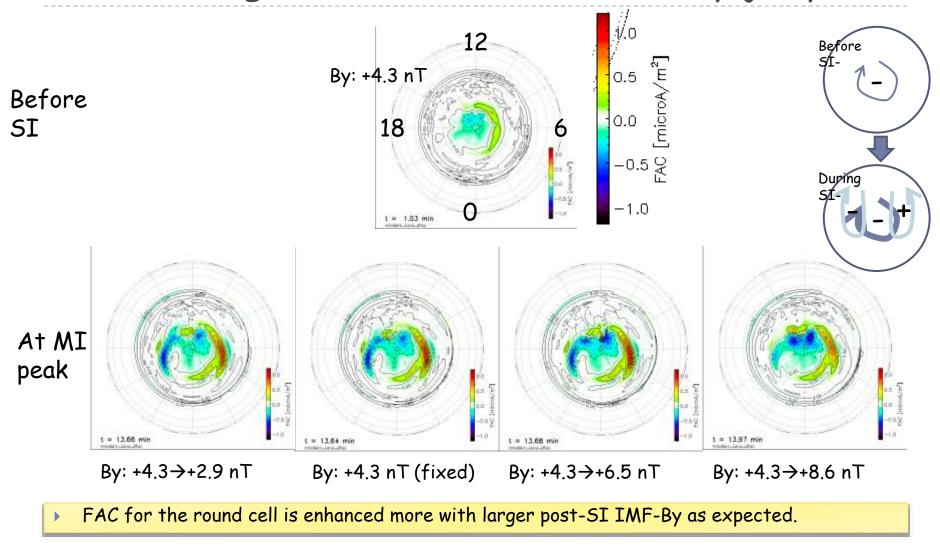
- Our hypothesis with the round cell evolution upon SI should be tested from a theoretical point of view.
- Objectives of the present (ongoing) work
  - We perform a set of global MHD simulation runs to reproduce the responses of the M-I coupled system to sudden changes of solar wind dynamic pressure.
  - The resultant profile of ionospheric potential/current and field-aligned current is compared with those observed by SD.

# 3. Data & Method Global MHD simulation of M-I coupled system

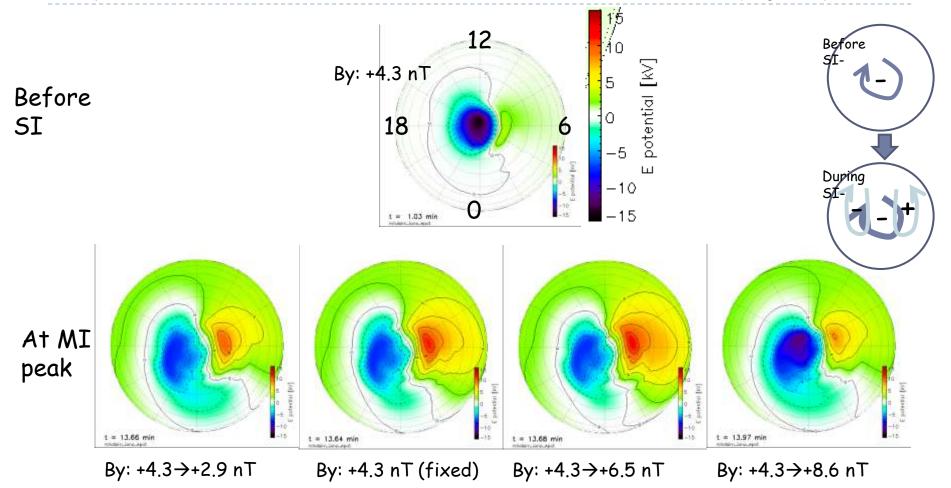


The same codes as those by Fujita+2012, Tanaka+2010.
Employed a realistic inhomogeneous ionospheric conductance.

## 4. Results FACs during SI+ for different IMF-By jumps



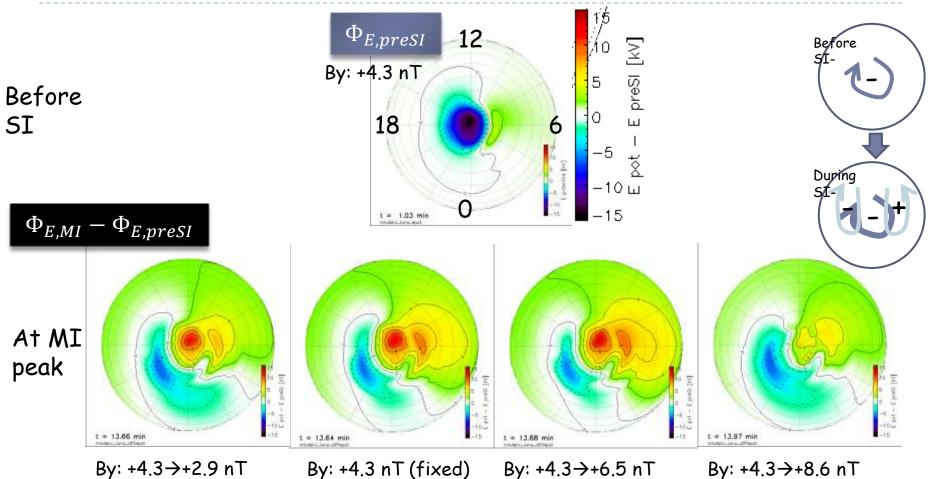
#### 4. Results E-potential during SI+ for different IMF-By jumps



> The expected superposition of round cell and SI cells is seen for all cases.

Hori, T., SI transient flow, SuperDARN meeting @STEL 9/14/2015

# 4. Results $\Delta E$ -potential during SI+ for different IMF-By jumps

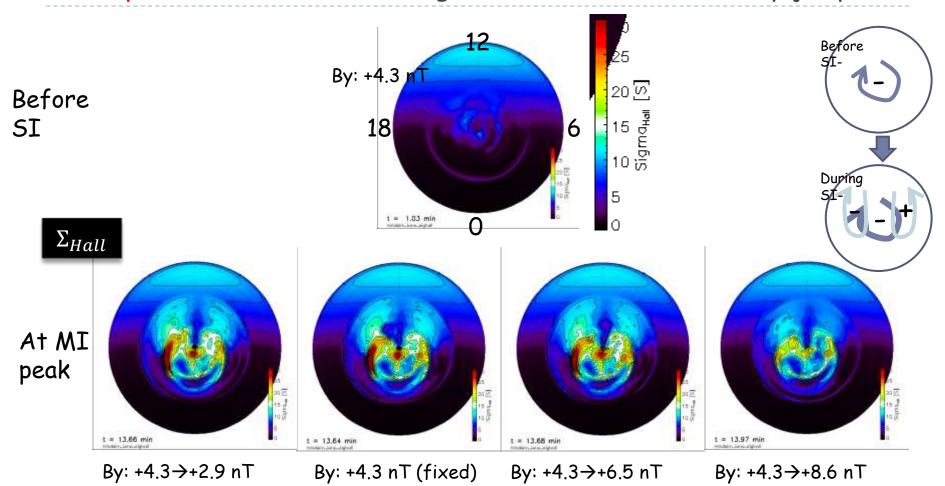


The side of enhanced anti-sunward flow is opposite to the observation. Why?

13

A residual positive round cell remains in polar cap → pre-SI negative round cell diminishes?

#### 5. Discussion **Ionospheric conductance** during **SI+** for different IMF-By jumps

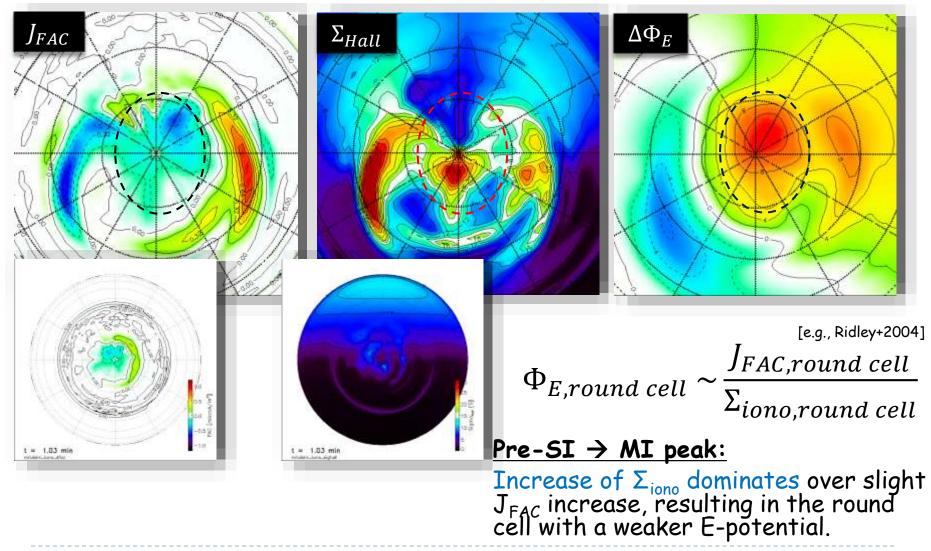


> The expected superposition of round cell and SI cells is seen for all cases.

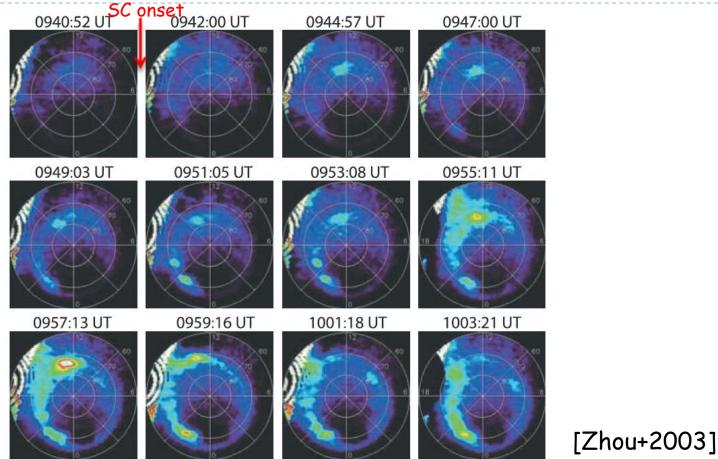
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# 5. Discussion Conductance, FAC, and E-potential relationship



# 5. Discussion Shock aurora profile by IMAGE/FUV



In reality, auroral luminosity does not enhance in the polar cap in the course of SC, suggesting no significant  $\Sigma_{iono}$  enhancement.

# 6. Summary & Conclusion

- The dawn-dusk asymmetric enhancement of transient sunward flow associated with SI- can be attributed to the round cell intensification with increasing |IMF-By| upon Psw drops. We have examined if the above features are reproduced by a global MHD simulation.
- FACs for the round cell do intensify upon arrival of larger IMF-By, consistent with our hypothesis.
- The resultant E-field potential, however, gives flow enhancement of opposite sense. This may be due to unrealistic increase/decrease of ionospheric conductance assumed in the simulation.