

SuperDARNで観測されたPc5帯ULF波動の モード解析

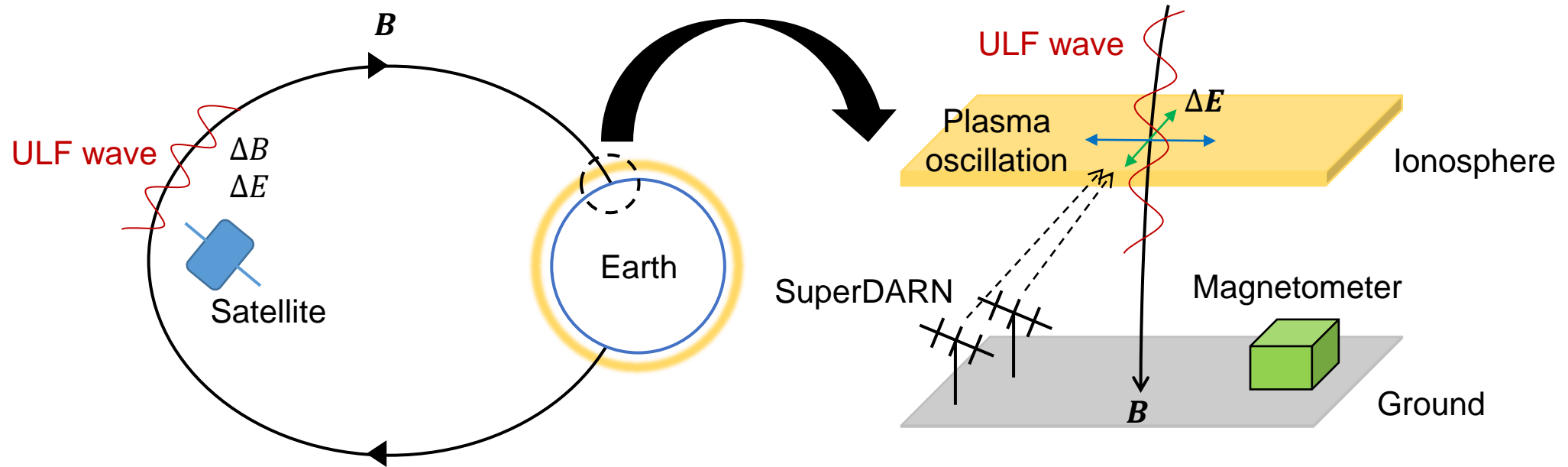
Mode analysis of ULF waves in the Pc5
frequency range observed by SuperDARN

Koki Morita, N. Nishitani, T. Hori

Institute for Space-Earth Environmental Research, Nagoya University (ISEE)



Introduction: ULF wave



◆ Ultralow Frequency (ULF) waves are mainly driven in the magnetosphere.

➤ Satellite data

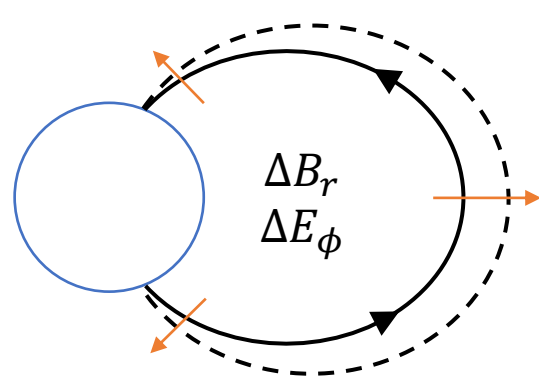


◆ ULF waves propagate along the earth's magnetic field lines and reach the ionosphere, causing perturbation of the ionospheric plasma motion and the magnetic field.

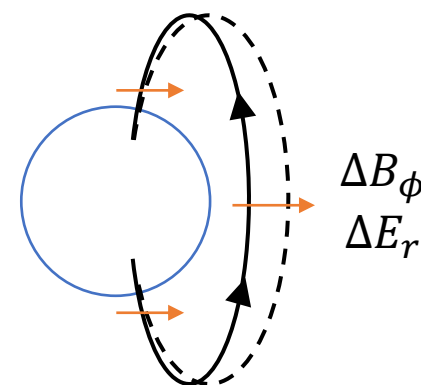
➤ SuperDARN / ground-based magnetometer data

Introduction: ULF wave modes

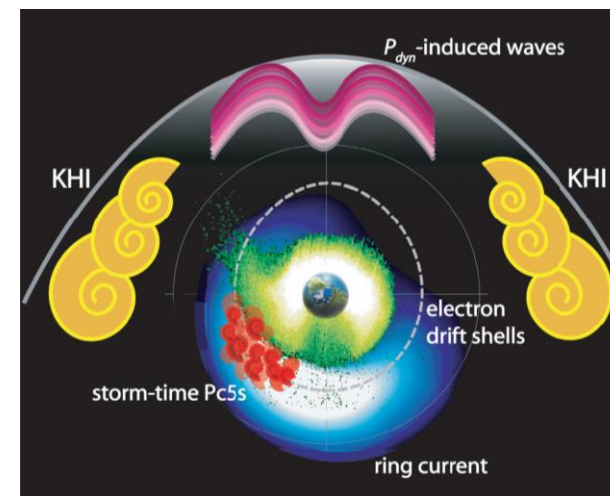
	Poloidal mode	Toroidal mode
Direction of ionospheric plasma oscillation	North-south perpendicular to the magnetic field lines	East-west
Primary driving source	<ul style="list-style-type: none"> ◆ Internal source <ul style="list-style-type: none"> • Drift mirror instability • Drift-bounce resonance instability ◆ External source <ul style="list-style-type: none"> • Solar wind dynamic pressure variations • Kelvin-Helmholtz Instability 	



Poloidal mode



Toroidal mode

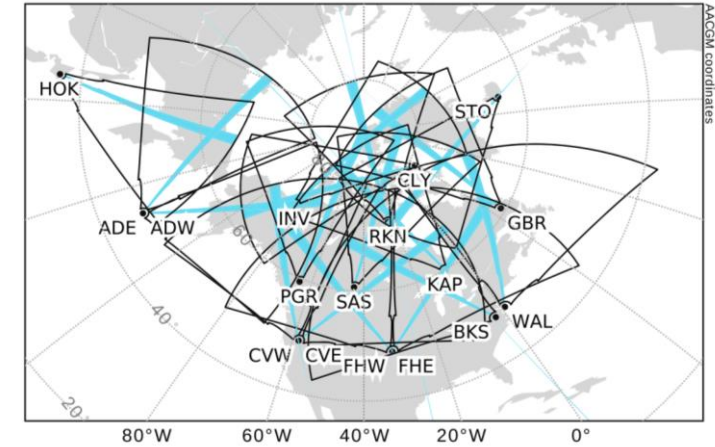


[Ukhorskiy et al., 2009]

Motivation

◆ Previous studies on ULF waves with SuperDARN

- Sakaguchi et al. (2012): MLAT 60-67°
 - The occurrence rate of Pc5 is higher on the night side than on the dayside.
- Shi et al. (2018): wide range of MLAT
 - This paper analyzed Pc5 waves with high time resolution data and showed an MLAT/MLT distribution.
- Wharton et al. (2019): MLAT 60-76°
 - This paper suggested a driving source with a wide frequency band because the frequencies seen by the radar and the magnetometer are different.



[Shi et al., 2018]

- These studies analyzed ULF waves in a single line-of-sight direction.
- Few studies addressed **occurrent characteristics depending on the wave mode.**

Purpose

◆ We analyze Pc5 waves with **multiple line-of-sight beam data**.

- Amplitude comparison → mode identification
- Phase comparison → m-number identification



◆ Statistical analysis on geomagnetic latitude (MLAT) and local time (MLT) dependence of the Pc5 wave modes and m-number

- Only the mode results are shown in this presentation.

Instrumentation and Data

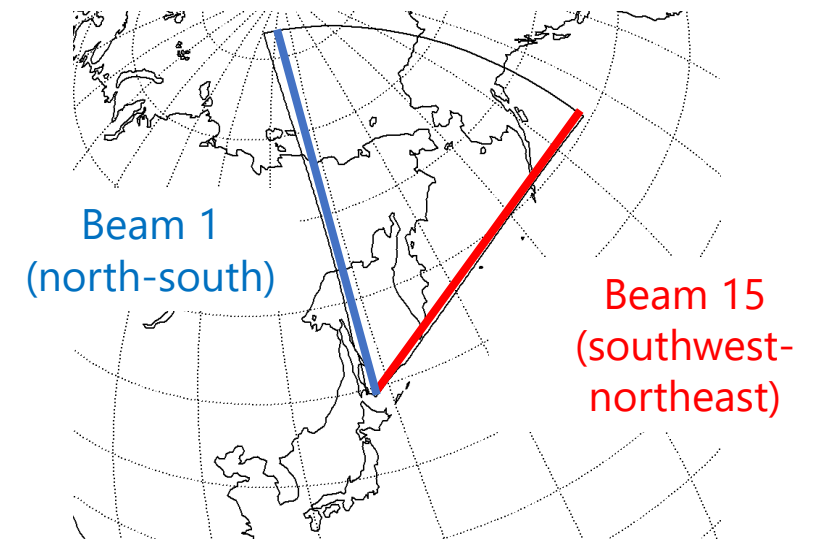
◆ SuperDARN Hokkaido East Radar (HOK)

- FOV is covered with 16 beams and a line-of-sight plasma velocity is measured for each beam direction.
- Magnetic latitude range of FOV: $\sim 40\text{-}80^\circ$



◆ We use **only ionospheric backscatter echoes** for this analysis.

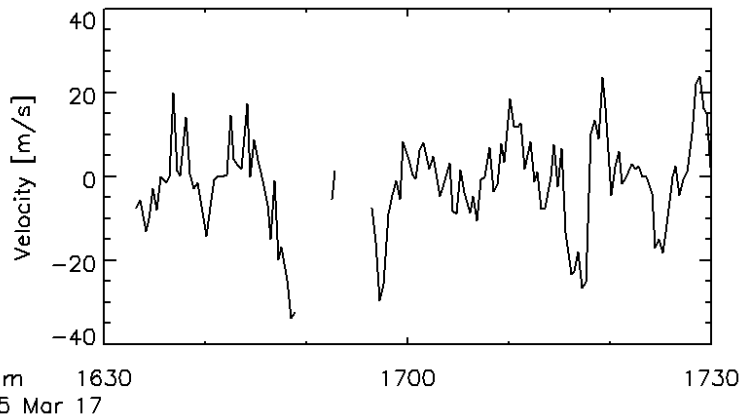
- Toroidal Pc5 waves cannot be found in ground/sea backscatter data.
- The results in this study are limited to the **nighttime**.



Method: Lomb-Scargle Periodogram

The Lomb-Scargle periodogram has a significant advantage in that it can easily be applied to **unevenly sampled data with data gaps**.

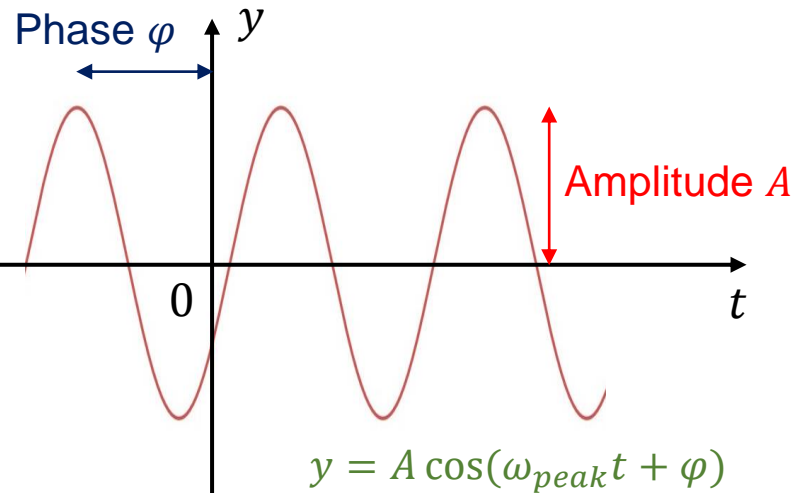
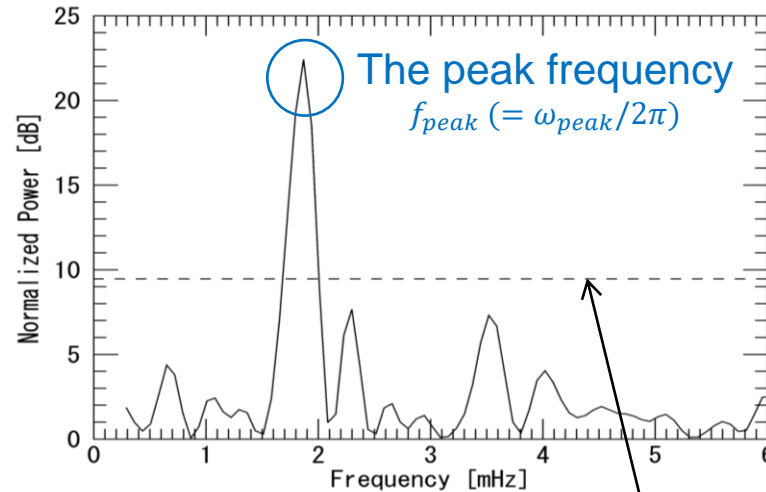
Median high-pass filtered
ionospheric Doppler velocity data



1-hr interval

incremented by 15 min iteratively

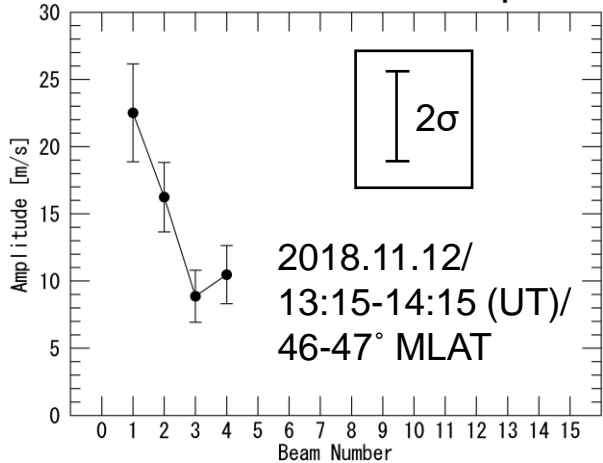
Lomb-Scargle Normalized Periodogram



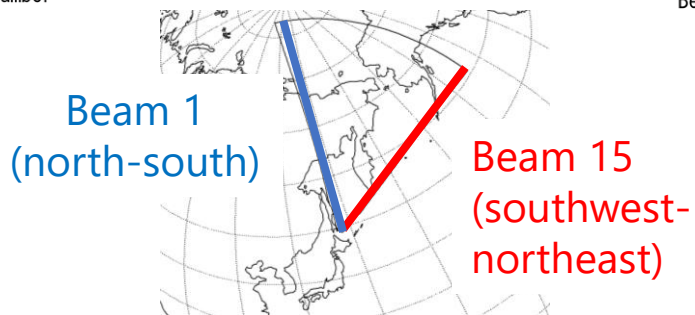
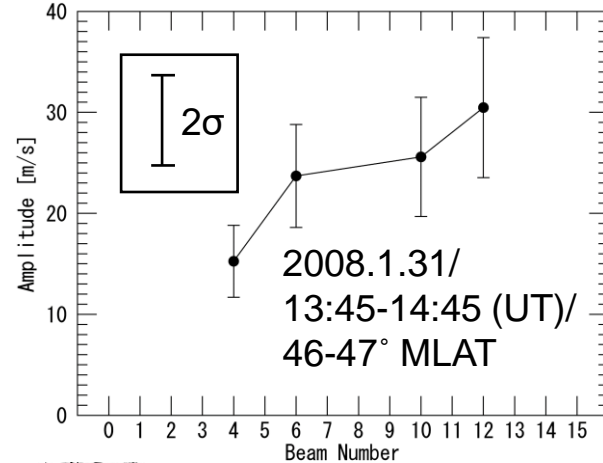
Method: Mode and M-number Identification

- ◆ We identify the mode of Pc5 by comparing the **amplitudes** at the peak frequency among different beams at a fixed MLAT (1° interval), assuming that the Pc5 wave has a constant amplitude with the same mode over the observed longitudinal range.

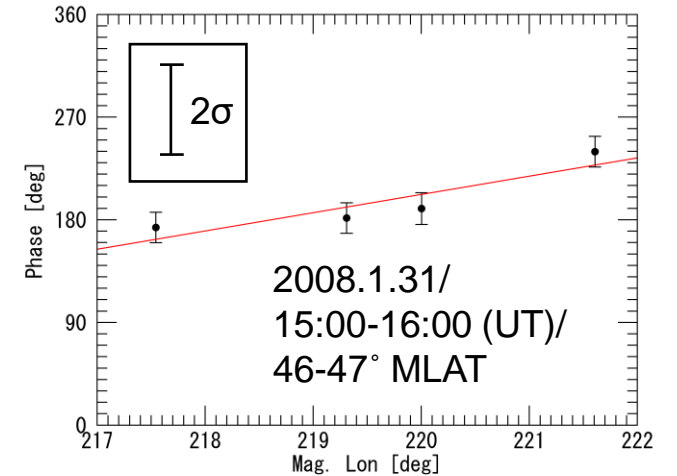
Poloidal mode wave amplitude



Toroidal mode wave amplitude



- ◆ We identify the m-number of Pc5 by comparing the **phases** among different geomagnetic longitudes at a fixed MLAT (1° interval).

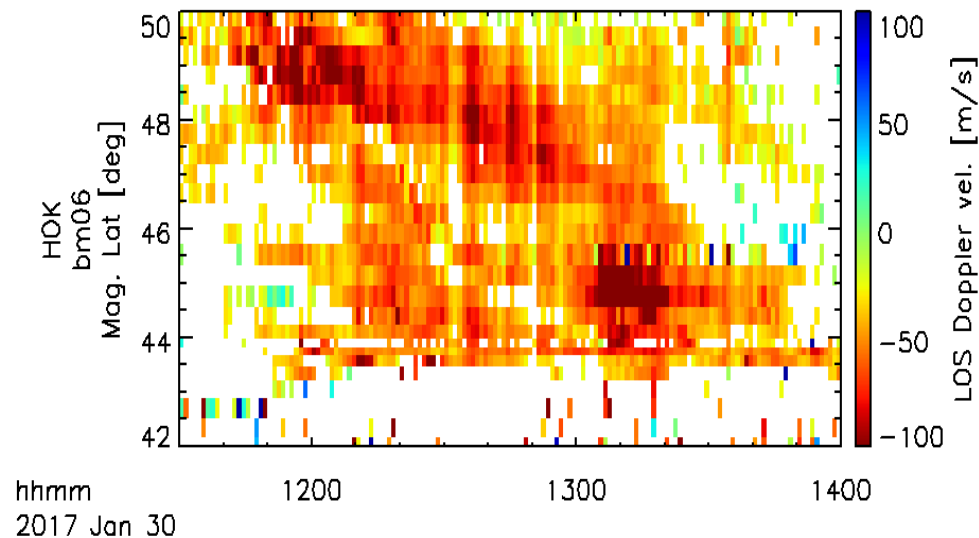


- $m \approx 16$
- Eastward propagation

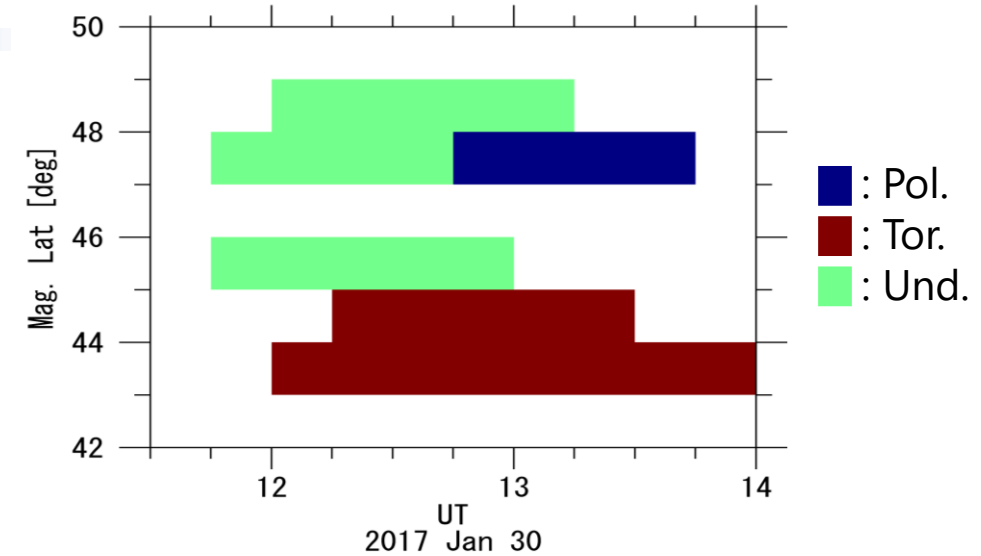
Case Results

- ◆ 2017.1.30
- ◆ The average peak frequency: ~2.0 mHz
- ◆ The results of the overlapping cells are average mode and m-number.
 - Poloidal & Toroidal → Undefined
 - Toroidal & Undefined → Toroidal

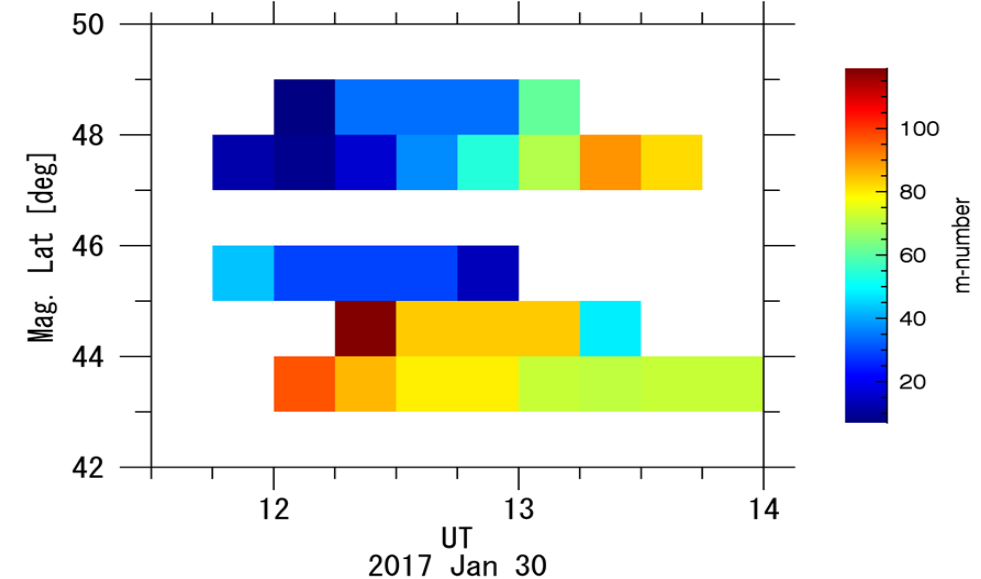
Range-Time-Intensity plot



The Pc5 wave mode results



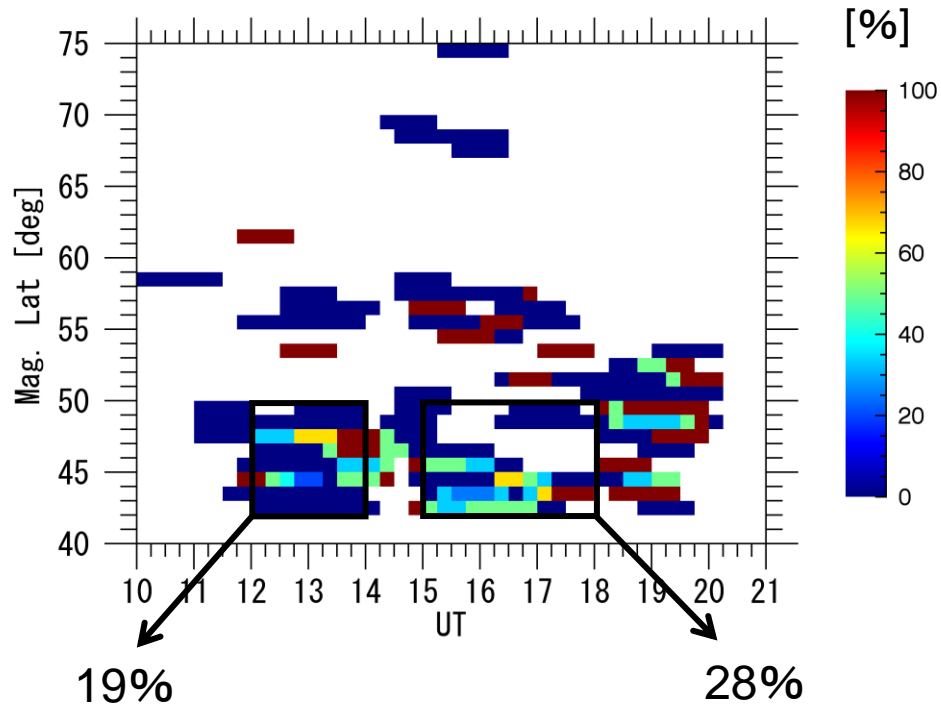
The Pc5 wave m-number results



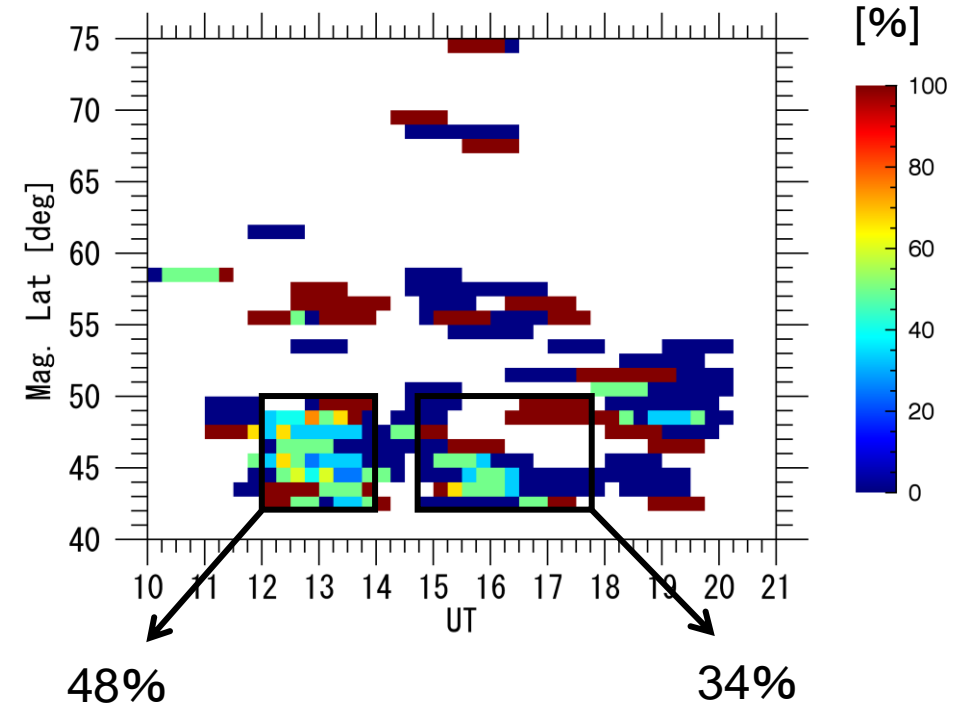
Statistical Results

◆ Analysis period: 2011-2020 (32 events)

Occurrence rates of **Poloidal** mode Pc5 waves
(only ionospheric backscatter echoes)



Occurrence rates of **Toroidal** mode Pc5 waves
(only ionospheric backscatter echoes)



MLT \approx UT + 9

The occurrence rate of toroidal mode Pc5 waves is higher than that of poloidal mode during 21-23 MLT and during 0-3 MLT in the range of 42-50° MLAT.

Discussion

Comparison with Liu et al. (2009)

- ◆ Liu et al. (2009) used electric and magnetic field measurements from the THEMIS satellite.
- ◆ From Figure 7 of Liu et al. (2009), we can see that the occurrence rate of toroidal mode Pc5 waves is higher than that of poloidal mode during 21-23 MLT and 0-3 MLT, in agreement with our results.
- ◆ Caution
 - Target magnetic latitude
 - Liu et al. (2009): $L > \sim 5$ (MLAT $> \sim 63.4^\circ$)
 - Our study: $42-50^\circ$ MLAT
 - Our statistical results in the target latitude region of Liu et al. (2009) have few data points.

Fig. 1

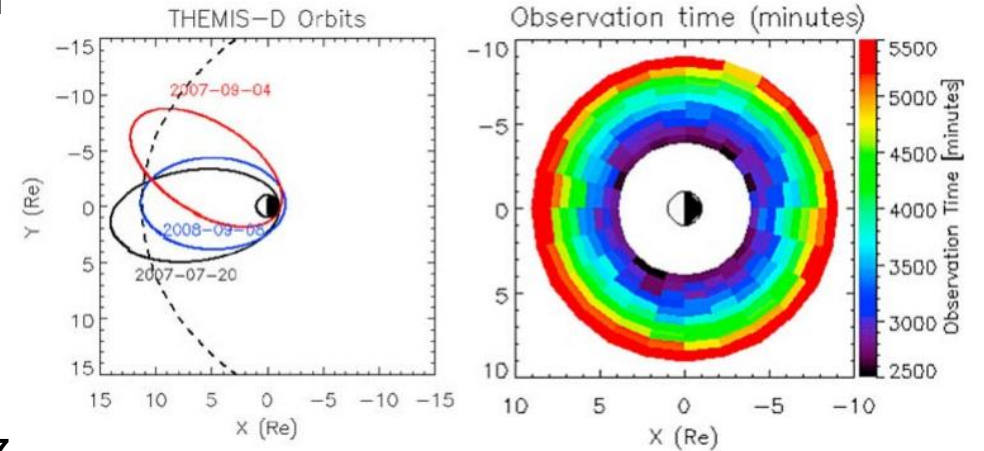
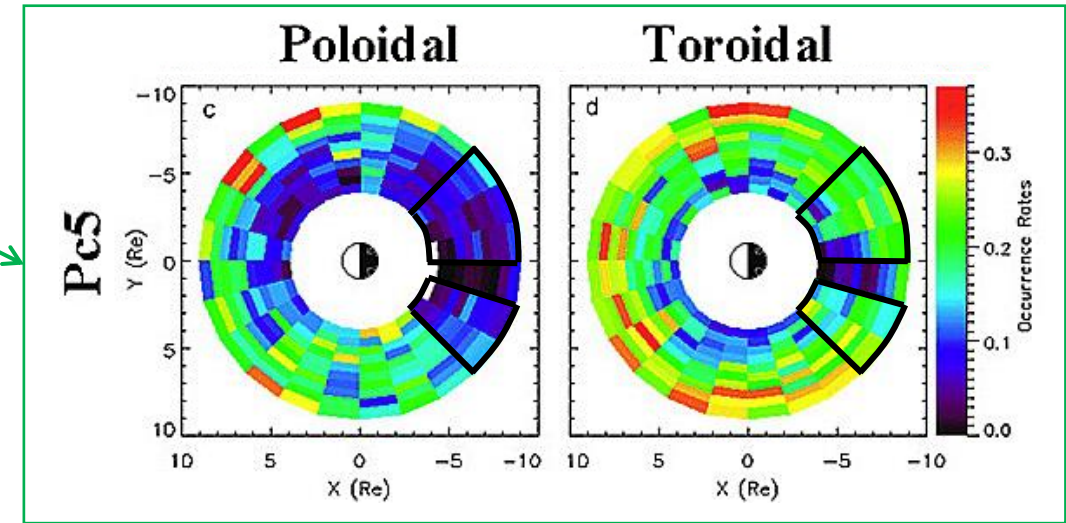


Fig. 7



[Liu et al., 2009]

Conclusions

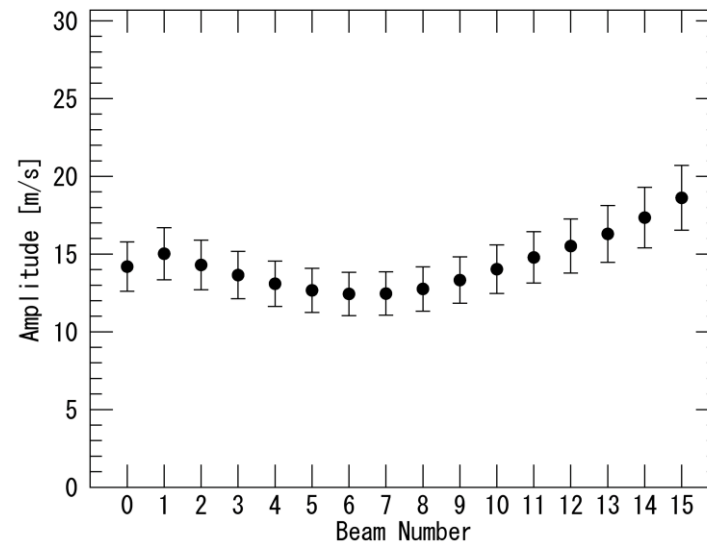
- ◆ The occurrence rate of toroidal mode Pc5 waves is higher than that of poloidal mode during 21-23 MLT and during 0-3 MLT in the range of 42-50° MLAT, in agreement with the statistical MLT results of Liu et al. (2009).
- ◆ To the best of our knowledge, it is the first study on the MLAT and MLT dependences of the Pc5 wave modes in the midlatitude region with SuperDARN.

Future Work

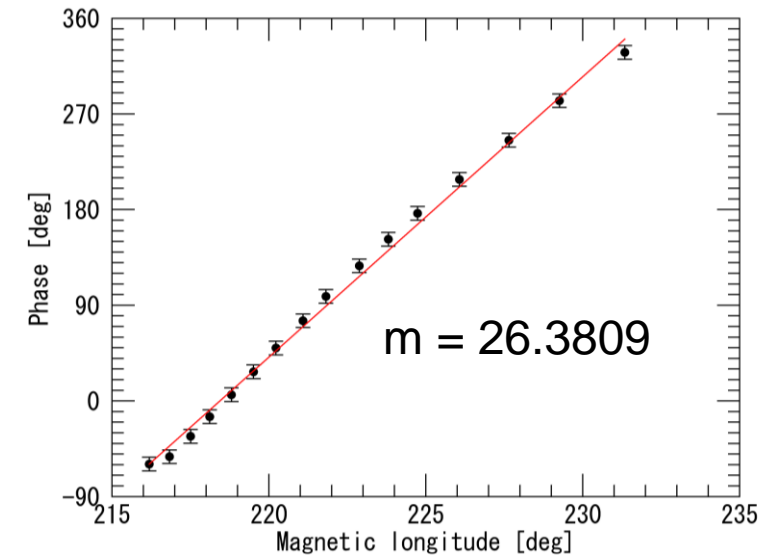
- ◆ We started simulating the amplitude and phase of the Doppler velocity at each beam using the modeled mode-coupled Pc5 waves with various parameters such as the poloidal/toroidal wave amplitude ratio and m-number.

Parameters

- Frequency: 2.0 mHz
- Poloidal wave amplitude: 15 m/s
- Toroidal wave amplitude: 30 m/s
- Phase difference: 120°
- M-number: 20
- Magnetic latitude: $45-46^\circ$



Undefined mode in this study



We discussed the error.

- ◆ We are evaluating our statistical analysis results by comparing with the model simulation results.